



On the interpretation  
of the loading  
correction of the  
aethalometer

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# On the interpretation of the loading correction of the aethalometer

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Received: 15 May 2015 – Accepted: 28 June 2015 – Published: 17 July 2015

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Published by Copernicus Publications on behalf of the European Geosciences Union.

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## Abstract

Aerosol optical properties were measured with a 7-wavelength aethalometer and a 3-wavelength nephelometer at the suburban site SORPES in Nanjing, China, in September 2013–January 2015. The aethalometer compensation parameter  $k$ , calculated with the Virkkula et al. (2007) method depended on the backscatter fraction, measured with the independent method, the integrating nephelometer. At  $\lambda = 660$  nm the daily-averaged compensation parameter  $k \approx 0.0017 \pm 0.0002$  and  $0.0042 \pm 0.0013$  when backscatter fraction at  $\lambda = 635$  nm was in the ranges of  $0.100 \pm 0.005$  and  $0.160 \pm 0.005$ , respectively. Also the wavelength dependency of the compensation parameter depended on the backscatter fraction: when  $b(\lambda = 525$  nm) was less than approximately 0.13 the compensation parameter decreased with wavelength and at larger  $b$  it increased with wavelength. This dependency has not been considered in any of the algorithms that are currently used for processing aethalometer data. The compensation parameter also depended on single-scattering albedo  $\omega_0$  so that  $k$  decreased with increasing  $\omega_0$ . For the green light ( $\lambda = 520$  nm) in the  $\omega_0$  range  $0.870 \pm 0.005$  the average ( $\pm$  standard deviation)  $k \approx 0.0047 \pm 0.006$  and in the  $\omega_0$  range  $0.960 \pm 0.005$   $k \approx 0.0028 \pm 0.0007$ . This difference was larger for the near-infrared light ( $\lambda = 880$  nm): in the  $\omega_0$  range  $0.860 \pm 0.005$   $k \approx 0.0055 \pm 0.0023$  and in the  $\omega_0$  range  $0.960 \pm 0.005$   $k \approx 0.0019 \pm 0.0011$ . The negative dependence of  $k$  on  $\omega_0$  was also shown with a simple theoretical analysis.

## 1 Introduction

Aerosols affect both local, regional, and global climate directly by scattering and absorbing solar radiation and indirectly by modifying cloud properties (e.g., IPCC, 2013). For the assessment of the direct radiative forcing it is crucial that both light scattering and absorption are measured accurately. Light scattering measurements with the nephelometer are well established but absorption is more difficult. An ideal method

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two aethalometers at two different flow rates and processed the data with a method that is a modified version of that in the dual-spot aethalometer.

Despite of the weaknesses, several authors have used the function  $f = 1 + kATN$  in post processing their aethalometer data and also filter samples analyzed with a reflectometer (e.g., Heal and Hammonds, 2014). It has been observed in many studies that the  $k$  varies in time and place. For example Park et al. (2010) found different values for the  $k$  in indoor and outdoor aerosol. Seasonal variation of it was presented already in the original paper: both in an urban and rural site the factor was higher in winter than in summer (Virkkula et al., 2007). No good explanation was given, it was just hypothesized that it was due to the variation of the single-scattering albedo which also has a seasonal cycle. A similar observation has been made also at other locations: for instance in East Rochester, New York, USA (Wang et al., 2011), at several sites in and around Beijing, China (Song et al., 2013) the  $k$  factors were larger in winter than in summer. Also Song et al. (2013) suggested this was probably due to darker aerosols. It is definitely expected that the compensation parameter depends on the darkness of the particles, since the more detailed algorithms to calculate  $\sigma_{ap}$  from the aethalometer data take the single-scattering albedo explicitly into account (e.g., Schmid et al., 2006; Collaud Coen et al., 2010).

Also the size of particles affects the absorption coefficients calculated from filter-based measurements. One of the reasons is that the penetration depth of the particles into the filter depends on their size and the depth affects the amount of light interactions with the particles (e.g., Arnott et al., 2005; Moteki et al., 2010; Nakayama et al., 2010). Lack et al. (2009) found that for particles larger than about 350 nm absorption measured with the Particle Soot Absorption Photometer (PSAP), another filter-based instrument, was significantly underestimated and concluded that the low bias of linked to the enhanced forward scattering from the larger particles. Müller et al. (2014) found that the asymmetry parameter – which is a function of the backscatter fraction – of the particles collected on the PSAP filter has significant effects on the derived  $\sigma_{ap}$ . It is

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humidity of the sample air is measured with the RH sensor of the nephelometer. In this study only those data were used during which RH was less than 50 %. When RH is higher particles grow significantly which affects all optical measurements. The World Meteorological Organization Global Atmosphere Watch (WMO/GAW) recommends for aerosol monitoring stations to keep sample air RH at  $45 \pm 5$  % (WMO, 2003).

### 2.1.1 Nephelometer

Total scattering coefficients ( $\sigma_{sp}$ ) and backscattering coefficients ( $\sigma_{bsp}$ ) at  $\lambda = 450, 525,$  and  $635$  nm were measured with an ECOTECH Aurora 3000 nephelometer. The scattering and backscattering coefficients are presented at STP conditions. The flow to the nephelometer was provided by the internal pump of the instrument. The averaging time was set to 5 min. The nephelometer was calibrated manually and zeros and spans were checked automatically according to the manual by using 1,1,1,2-Tetrafluoroethane (R-134) as the calibration gas.

The raw total scattering coefficients were corrected for truncation errors by calculating first the Ångström exponents from the non-corrected scattering coefficients and then following the formulas presented by Müller et al. (2011) where the tabulated factors for no cutoff at the inlet were used. To be used in the aethalometer data processing the truncation-corrected  $\sigma_{sp}$  at the nephelometer wavelengths were interpolated and extrapolated to the aethalometer wavelengths assuming that the Ångström exponent of scattering was constant over the wavelength range.

The backscatter fractions ( $b = \sigma_{bsp}/\sigma_{sp}$ ) were calculated as the ratio of the backscattering coefficient and the truncation-corrected total scattering coefficients at the nephelometer wavelengths. The backscattering coefficients were not interpolated or extrapolated.

## 2.1.2 Aethalometer

A 7-wavelength aethalometer (AE-31) was used for measuring light absorption at  $\lambda = 370, 470, 520, 590, 660, 880, \text{ and } 950 \text{ nm}$ . The aethalometer reports BC concentrations but from these data absorption coefficients were calculated as will be discussed below. The flow was provided by the internal pump, it was set to 5 LPM at  $t = 20^\circ\text{C}$  and  $p = 1013 \text{ mbar}$ . Flow checks with a Gilibrator flow meter showed that the flow was  $4.7 \pm 0.2 \text{ LPM}$  at the same conditions. Concentrations were converted to STP (Standard Temperature and Pressure, 273.15 K, 1013.25 mbar), taking the flow calibrations into account. The filter spots were set to change when the maximum attenuation (ATN) exceeded 125. The average and standard deviation of the last ATN values before filter spot changes were  $127 \pm 3, 99 \pm 6, 87 \pm 7, 79 \pm 7, 73 \pm 8, 54 \pm 7, \text{ and } 49 \pm 7$  for  $\lambda = 370, 470, 520, 590, 660, 880, \text{ and } 950 \text{ nm}$ , respectively. These are given here to be used for evaluating the effect of the correction function.

## 2.2 Calculation of the compensation parameter

The core of the present paper is to analyze factors affecting the compensation parameter  $k$  that is used to correct BC concentrations in

$$BC_{\text{corr}} = (1 + k \cdot \text{ATN})BC_0, \quad (1)$$

where  $BC_0$  is the original non-corrected BC concentration and ATN is the attenuation reported by the aethalometer. The  $k$  of filter spot  $i$  was calculated from

$$k = \frac{1}{\text{ATN}_{i, \text{last}}} \left( \frac{BC_{0, i+1, \text{first}}}{BC_{0, i, \text{last}}} - 1 \right), \quad (2)$$

where  $\text{ATN}_{i, \text{last}}$  is the last attenuation of filter spot  $i$  before the filter spot change,  $BC_{0, i, \text{last}}$  and  $BC_{0, i+1, \text{first}}$  are the original non-corrected BC concentrations of the last

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measurement point of filter spot  $i$  and the first measurement point of spot  $i+1$ , respectively. In practice the averages of three last measurement points of filter spot  $i$  and the averages of three first measurements of filter spot  $i+1$  were used, as in the original paper (Virkkula et al., 2007). At this point it is worth noting the analogy of the  $k$  factor in Eq. (1) and that of  $BC_{\text{corr}} = BC_0 / (1 - k \cdot \text{ATN})$  which is used in the dual-spot aethalometer, model AE33 (Drinovec et al., 2015).  $1 / (1 - k \text{ATN})$  is the sum of a geometric series  $1 + k \text{ATN} + (k \text{ATN})^2 + \dots$ . Typically published values of  $k$  are less than 0.01 and aethalometers are usually changing spots when ATN is less than 100, so the terms  $(k \text{ATN})^n$  with  $n > 1$  are small and  $1 / (1 - k \text{ATN}) \approx 1 + k \text{ATN}$ . This suggests that the results to be shown below are qualitatively valid also for the new model.

### 2.3 Calculation of absorption coefficients

The aethalometer data were first used to calculate the raw, uncorrected absorption coefficients, here  $\sigma_0$ , by multiplying the the original non-corrected BC concentration ( $BC_0$  above) given by the aethalometer with the wavelength-dependent BC mass absorption coefficient used by the instrument's software. To calculate the absorption coefficients ( $\sigma_{\text{ap}}$ ) several algorithms have been presented which in principle can be expressed in the form of

$$\sigma_{\text{ap}} = \frac{f \sigma_0 - s \sigma_{\text{sp}}}{C_{\text{ref}}}, \quad (3)$$

where  $f$  is a loading correction function,  $s$  is a fraction of light scattering coefficient  $\sigma_{\text{sp}}$  that causes reduction of light transmittance and would be interpreted as absorption (= apparent absorption) if not taken into account, and  $C_{\text{ref}}$  the multiple scattering correction factor. Note, however, that  $s$  is not any constant factor but also a function (e.g., Arnott et al., 2005; Collaud Coen et al., 2010). In the present work absorption coefficients were calculated according to both Arnott et al. (2005) and Collaud Coen et al. (2010) algorithms with the respective mean  $C_{\text{ref}}$  values of 4.12 and 4.26 obtained for the Cabauw station by Collaud Coen et al. (2010). The differences of the absorption







dependency (Fig. 2a). A line

$$k = a_k \lambda + k_0 \quad (4)$$

was fit through the 7  $k$  values obtained for each filter spot change of the whole data set. Only the slope  $a_k$  is of interest here. Its interpretation is simple: when  $a_k > 0$  the compensation parameters increase with wavelength, when  $a_k < 0$ , the compensation parameters decrease with wavelength. As noted already earlier, the compensation parameters obtained for individual spot changes are noisy. Therefore more relevant information was obtained when the compensation parameters from all spot changes were classified according to the associated filter-spot-averaged backscatter fraction of green light and simple descriptive statistics were calculated. Figure 2b shows the averages, medians and the 25th to 75th percentile ranges of the cumulative distributions of the compensation parameters at three different backscatter fraction ranges. The lines shown in the figure were fit to the average compensation parameters in each wavelength and bin of  $b$ . Note again that  $\omega_0$  was high when  $b$  was low and low when  $b$  was high and  $a_k$  increased with increasing  $b$ .

Another interesting observation can be made on Fig. 2: the range of compensation parameters is the larger the longer the wavelength is. This suggests that the longer wavelengths are more sensitive to the factors affecting the compensation. The near-infrared wavelength at  $\lambda = 880$  nm is the one that is used in most aethalometers, even single-wavelength ones, and therefore more attention will be paid to it than to the longest wavelength ( $\lambda = 950$  nm).

A shorter, two-month time series of the data is presented in Fig. 3. In addition to those quantities presented in Fig. 1 also the slope  $a_k$  calculated for each filter spot change and the respective daily averages are shown (Fig. 3d). There are some interesting features in the time series of  $\sigma_{\text{sp}}$ ,  $\sigma_{\text{ap}}$ ,  $\omega_0$  and  $b$ . First, the most polluted episodes, for instance on 4–8 December were associated with the highest  $\sigma_{\text{sp}}$  and the highest  $\omega_0$ , suggesting that the contribution of light-absorbing material, mainly BC, to the aerosol mass in the highly-polluted air was clearly lower than during the less-polluted periods,

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compensation parameters, for reasons discussed already above. There is a positive correlation, however, and it is demonstrated visually by classifying the daily-averaged compensation parameters at four wavelengths ( $\lambda = 470, 520, 660, \text{ and } 880 \text{ nm}$ ) into bins of backscatter fractions at the nearest nephelometer wavelengths ( $\lambda = 450, 525, \text{ and } 635 \text{ nm}$ ) (Fig. 6). The width of the backscatter fraction bins was 0.005. The averages, 5th, 25th, 50th, 75th, and 95th percentiles of the cumulative distribution of the compensation parameters in each bin were calculated.

The compensation parameter medians and averages correlated positively with the backscatter fractions, and so did the other percentiles but their correlation was weaker. Note that the slopes of the linear regressions of  $k$  vs.  $b$  are almost the same when the wavelength of both  $k$  and  $b$  are approximately the same (Fig. 6a–c). When  $k$  at  $\lambda = 880 \text{ nm}$  is plotted against  $b$  at  $\lambda = 635 \text{ nm}$ , the slope is almost twice as high.

Instead of paying much attention to the  $R^2$  values, it is more relevant to test the statistical significance of the slope of the regression of compensation parameter vs. backscatter fraction, i.e.  $\beta_1$  in  $k = \beta_1 b + \beta_0$ . The null hypothesis that the slope is not dependent on the  $b$ , i.e.,  $\beta_1 = 0$  was tested using test statistics given by the estimate of the slope divided by its standard error ( $t = \beta_1 / \text{s.e.}$ ). The test statistics were compared with the Student's  $t$  distribution on  $n - 2$  (sample size – number of regression coefficients) degrees of freedom. The regressions were calculated both for each individual filter change and for daily averages. Four compensation parameter–backscatter fraction pairs were used: blue:  $k(\lambda = 470 \text{ nm})$  vs.  $b(\lambda = 450 \text{ nm})$ , green:  $k(\lambda = 520 \text{ nm})$  vs.  $b(\lambda = 525 \text{ nm})$ , red:  $k(\lambda = 660 \text{ nm})$  vs.  $b(\lambda = 635 \text{ nm})$ , and red-near-infrared:  $k(\lambda = 880 \text{ nm})$  vs.  $b(\lambda = 635 \text{ nm})$ . The last combination differs from the other three in that the wavelength of  $k$  and  $b$  are not close to the same like in the other cases. The reason this is considered to be relevant here is that most aethalometers have the 880 nm wavelength and in most 3-wavelength nephelometers the longest wavelength is 600–700 nm. The results are presented in Table 1. The  $p$  values for all wavelengths are all low ( $< 0.001$ ) which gives strong evidence against the null hypothesis, indicating that the slope is not 0 and that there is a linear relationship between  $k$  and  $b$ .



the correction is the larger the larger the wavelength is and for large particles the other way round.

### 3.4 Effect of single-scattering albedo

The relationship of the single-scattering albedo and the compensation parameter was analyzed analogically. The  $k$ 's were classified into bins of  $\omega_0$  at four aethalometer wavelengths ( $\lambda = 470, 520, 660, \text{ and } 880 \text{ nm}$ ). The width of the  $\omega_0$  bins was 0.01. The averages, 5th, 25th, 50th, 75th, and 95th percentiles of the cumulative distribution of the compensation parameters were calculated. The bin averages and medians decreased almost monotonically with increasing  $\omega_0$ , but the ranges were large (Fig. 8). Note that in Fig. 8 the  $k$  vs.  $\omega_0$  relationship of the bin averages is plotted both by using the Colaud Coen et al. (2010) algorithm and the Arnott et al. (2005) algorithm simply to show that the main relationship: decreasing  $k$  with increasing  $\omega_0$  did not depend on the algorithm used for calculating absorption coefficients  $\sigma_{\text{ap}}$ . It is worth noting at this point that the absolute values of  $\sigma_{\text{ap}}$  and  $\omega_0$  are very uncertain because of the uncertainty of the multiple scattering correction factor  $C_{\text{ref}}$ . The  $\omega_0$  values shown in Fig. 8 were calculated with  $C_{\text{ref}}$  values of 4.12 and 4.26 as explained above but if  $C_{\text{ref}}$  is smaller  $\omega_0$  is lower than that shown in Fig. 8. This would not change the main result:  $k$  decreases with increasing  $\omega_0$ . The decrease is also statistically significant. Linear regression of  $k$  vs.  $\omega_0$  was calculated both for individual filter changes and for daily averages, as above for  $k$  vs.  $b$ . The statistics are presented in Table 2. The  $p$  values are somewhat higher than in Table 1 but still low enough to conclude that the relationship is statistically significant.

A simple theoretical explanation for the decreasing  $k$  with increasing  $\omega_0$  can be given. If it is assumed that (1) the loading correction function  $f$  in Eq. (3) equals  $1 + k\text{ATN}$  and (2) that the dependence on scattering coefficient is incorporated in the compensation parameter the equation for absorption coefficient becomes

$$\sigma_{\text{ap}} = \frac{1 + k\text{ATN}}{C_{\text{ref}}} \sigma_0. \quad (5)$$

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On the other hand, if it is assumed that absorption coefficient is calculated from Eq. (3) where  $f$  is not the same function as in Eq. (5) and where the dependence on scattering coefficient is explicitly presented and if the two Eqs. (3) and (5) are set equal the compensation parameter can be solved as

$$\begin{aligned} \frac{1 + kATN}{C_{\text{ref}}}\sigma_0 &= \frac{f\sigma_0 - S\sigma_{\text{sp}}}{C_{\text{ref}}} \\ \Leftrightarrow k &= \frac{1}{ATN} \left( f - 1 - \frac{S\sigma_{\text{sp}}}{\sigma_0} \right). \end{aligned} \quad (6)$$

When Eq. (3) is again rearranged as  $\sigma_0 = (C_{\text{ref}}\sigma_{\text{ap}} + S\sigma_{\text{sp}})/f$  and inserted in Eq. (6) the compensation parameter can be expressed as

$$k = \frac{1}{ATN} \left( f - 1 - \frac{S\sigma_{\text{sp}}}{\frac{C_{\text{ref}}\sigma_{\text{ap}} + S\sigma_{\text{sp}}}{f}} \right) = \frac{1}{ATN} \left( f \left( 1 - \frac{\sigma_{\text{sp}}}{\frac{C_{\text{ref}}}{S}\sigma_{\text{ap}} + \sigma_{\text{sp}}} \right) - 1 \right). \quad (7)$$

The term  $\sigma_{\text{sp}} / \left( \frac{C_{\text{ref}}}{S}\sigma_{\text{ap}} + \sigma_{\text{sp}} \right)$  is not exactly identical to single-scattering albedo  $\omega_0$ , but also it approaches unity when  $\omega_0$  approaches unity and  $f(1 - \sigma_{\text{sp}} / (\frac{C_{\text{ref}}}{S}\sigma_{\text{ap}} + \sigma_{\text{sp}})) \rightarrow 0$  and then  $k$  may even become negative. In other words, the compensation parameter can be negative but the resulting absorption coefficient can never be negative, this sets the limit to it.

The relationship of  $a_k$  and single-scattering albedo is very similar to that of  $k$  and  $\omega_0$ : also  $a_k$  decreases with increasing  $\omega_0$  (Fig. 9). In other words for darker aerosols,  $\omega_0$  less than approximately 0.92 the correction increases with wavelength and at higher  $\omega_0$  it decreases with wavelength. No theoretical explanation could be given at this point.

### 3.5 Separating the effects of single-scattering albedo and backscatter fraction

The above analysis showed that the compensation parameter depends both on the single-scattering albedo and the backscatter fraction. Which of them is a more dom-

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don't show this, it should not be considered proven. To study it in a laboratory, one should produce aerosols that have the same size (and  $b$ ) but variable  $\omega_0$ . This is difficult because if variations in  $\omega_0$  are accomplished by coating pure BC by condensing some scattering material, the particles grow and  $b$  decreases.

## 4 Conclusions

Aerosol optical properties were measured with a 7-wavelength aethalometer and a 3-wavelength nephelometer at the suburban site SORPES in Nanjing, China, in September 2013–January 2015. The most important result obtained from the analysis of the data is that quantities calculated from two independent methods, i.e., backscattering fraction measured with the nephelometer and the compensation parameter  $k$ , calculated from the aethalometer data with the Virkkula et al. (2007) algorithm were correlated. At  $\lambda = 660$  nm the daily-averaged compensation parameter  $k \approx 0.0017 \pm 0.0002$  and  $0.0042 \pm 0.0013$  when backscatter fraction at  $\lambda = 635$  nm was in the ranges of  $0.100 \pm 0.005$  and  $0.160 \pm 0.005$ , respectively. Also the wavelength dependency of the compensation parameter depended on the backscatter fraction: when  $b(\lambda = 525$  nm) was less than approximately 0.13 the compensation parameter decreased with wavelength and at larger  $b$  it increased with wavelength. This dependency has not been considered in any of the algorithms that are currently used for processing aethalometer data. The compensation parameter also depended on single-scattering albedo  $\omega_0$  so that  $k$  decreased with increasing  $\omega_0$ . For the green light ( $\lambda = 520$  nm) in the  $\omega_0$  range  $0.870 \pm 0.005$  the average ( $\pm$  standard deviation)  $k \approx 0.0047 \pm 0.006$  and in the  $\omega_0$  range  $0.960 \pm 0.005$   $k \approx 0.0028 \pm 0.0007$ . This difference was larger for the near-infrared light ( $\lambda = 880$  nm): in the  $\omega_0$  range  $0.860 \pm 0.005$   $k \approx 0.0055 \pm 0.0023$  and in the  $\omega_0$  range  $0.960 \pm 0.005$   $k \approx 0.0019 \pm 0.0011$ . The negative dependence on  $\omega_0$  was also shown with a simple theoretical analysis. The interpretation of the above results is complicated by the fact that  $b$  and  $\omega_0$  were not independent: the darkest aerosols had the highest backscatter fractions. The selection of one narrow  $\omega_0$  bin and classifying

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and the other way round when comparing with the single-scattering albedo. This can be considered as a recommendation for future research.

*Acknowledgements.* The research was supported by the Jiangsu Provincial Natural Science Fund (No.BK20 140 021), National Science Foundation of China (D051/41275129) and Academy of Finland's Centre of Excellence program (Centre of Excellence in Atmospheric Science – From Molecular and Biological processes to The Global Climate, project no. 272041).

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**Table 1.** Regression statistics ( $y = \beta_1 x + \beta_0$ ) of compensation parameter vs. backscatter fraction. s.e.: standard error of  $\beta_1$ ; 95 % confidence range of  $\beta_1$ ; d.f.: degrees of freedom;  $t = \beta_1/\text{s.e.}$ ;  $p$ :  $p$  value of the Student's  $t$  distribution.

| Calculated by using each filter change |          |      |           |       |                       |      |      |                       |  |
|--|----------|------|-----------|-------|-----------------------|------|------|-----------------------|--|
| $x$                                    | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$  | $p$                   |  |
| $b(450)$                               | $k(470)$ | 0.18 | 0.024     | 0.003 | (0.018 – 0.030)       | 2048 | 8.2  | $5.6 \times 10^{-16}$ |  |
| $b(525)$                               | $k(520)$ | 0.22 | 0.036     | 0.004 | (0.029 – 0.043)       | 2048 | 10.3 | $2.9 \times 10^{-24}$ |  |
| $b(635)$                               | $k(660)$ | 0.21 | 0.032     | 0.003 | (0.026 – 0.039)       | 2048 | 9.5  | $5.7 \times 10^{-21}$ |  |
| $b(635)$                               | $k(880)$ | 0.26 | 0.056     | 0.005 | (0.047 – 0.065)       | 2048 | 12.4 | $4.3 \times 10^{-34}$ |  |
| Calculated by using all daily averages |          |      |           |       |                       |      |      |                       |  |
| $x$                                    | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$  | $p$                   |  |
| $b(450)$                               | $k(470)$ | 0.28 | 0.017     | 0.003 | (0.011 – 0.024)       | 320  | 5.2  | $3.0 \times 10^{-07}$ |  |
| $b(525)$                               | $k(520)$ | 0.38 | 0.029     | 0.004 | (0.021 – 0.037)       | 320  | 7.3  | $2.3 \times 10^{-12}$ |  |
| $b(635)$                               | $k(660)$ | 0.32 | 0.025     | 0.004 | (0.017 – 0.033)       | 320  | 6.   | $6.5 \times 10^{-09}$ |  |
| $b(635)$                               | $k(880)$ | 0.43 | 0.048     | 0.006 | (0.037 – 0.060)       | 320  | 8.4  | $1.3 \times 10^{-15}$ |  |

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**Table 2.** Regression statistics ( $y = \beta_1 x + \beta_0$ ) of compensation parameter vs. single-scattering albedo. Detailed column description as in Table 1.

| Calculated by using each filter change |          |      |           |       |                      |   |         |      |      |                       |
|--|----------|------|-----------|-------|----------------------|---|---------|------|------|-----------------------|
| $x$                                    | $y$      | $r$  | $\beta_1$ | s.e.  | 95% confidence range |   | d.f.    | $t$  | $p$  |                       |
| $\omega_0(470)$                        | $k(470)$ | 0.07 | -0.006    | 0.002 | (-0.009              | - | -0.002) | 2048 | -3.4 | $7.3 \times 10^{-04}$ |
| $\omega_0(520)$                        | $k(520)$ | 0.08 | -0.008    | 0.002 | (-0.012              | - | -0.004) | 2048 | -3.8 | $1.3 \times 10^{-04}$ |
| $\omega_0(660)$                        | $k(660)$ | 0.07 | -0.007    | 0.002 | (-0.012              | - | -0.003) | 2048 | -3.1 | $2.0 \times 10^{-03}$ |
| $\omega_0(880)$                        | $k(880)$ | 0.08 | -0.010    | 0.003 | (-0.016              | - | -0.004) | 2048 | -3.5 | $4.7 \times 10^{-04}$ |
| Calculated by using all daily averages |          |      |           |       |                      |   |         |      |      |                       |
| $x$                                    | $y$      | $r$  | $\beta_1$ | s.e.  | 95% confidence range |   | d.f.    | $t$  | $p$  |                       |
| $\omega_0(470)$                        | $k(470)$ | 0.27 | -0.011    | 0.002 | (-0.016              | - | -0.007) | 320  | -5.1 | $5.9 \times 10^{-07}$ |
| $\omega_0(520)$                        | $k(520)$ | 0.31 | -0.015    | 0.003 | (-0.020              | - | -0.010) | 320  | -5.9 | $9.6 \times 10^{-09}$ |
| $\omega_0(660)$                        | $k(660)$ | 0.26 | -0.015    | 0.003 | (-0.021              | - | -0.009) | 320  | -4.8 | $2.0 \times 10^{-06}$ |
| $\omega_0(880)$                        | $k(880)$ | 0.25 | -0.019    | 0.004 | (-0.027              | - | -0.011) | 320  | -4.5 | $8.1 \times 10^{-06}$ |

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**Table 3.** Regression statistics ( $y = \beta_1 x + \beta_0$ ) of compensation parameter vs. backscatter fraction at a limited single-scattering albedo range. Detailed column description as in Table 1.

| Calculated by using those filter changes during which $\omega_0(520\text{ nm}) = 0.930 \pm 0.005$ |          |      |           |       |                       |      |     |                       |  |
|---|----------|------|-----------|-------|-----------------------|------|-----|-----------------------|--|
| $x$   | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$ | $p$                   |  |
| $b(450)$  | $k(470)$ | 0.24 | 0.051     | 0.012 | (0.027 – 0.075)       | 290  | 4.2 | $3.5 \times 10^{-05}$ |  |
| $b(525)$  | $k(520)$ | 0.31 | 0.065     | 0.012 | (0.042 – 0.089)       | 290  | 5.6 | $5.8 \times 10^{-08}$ |  |
| $b(635)$  | $k(660)$ | 0.29 | 0.059     | 0.011 | (0.037 – 0.081)       | 290  | 5.3 | $2.8 \times 10^{-07}$ |  |
| $b(635)$  | $k(880)$ | 0.36 | 0.101     | 0.015 | (0.071 – 0.131)       | 290  | 6.7 | $1.4 \times 10^{-10}$ |  |
| Calculated by using those daily averages during which $\omega_0(520\text{ nm}) = 0.930 \pm 0.005$ |          |      |           |       |                       |      |     |                       |  |
| $x$   | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$ | $p$                   |  |
| $b(450)$  | $k(470)$ | 0.16 | 0.013     | 0.012 | (–0.011 – 0.038)      | 48   | 1.1 | 0.28                  |  |
| $b(525)$  | $k(520)$ | 0.21 | 0.017     | 0.012 | (–0.007 – 0.042)      | 48   | 1.5 | 0.15                  |  |
| $b(635)$  | $k(660)$ | 0.21 | 0.017     | 0.011 | (–0.005 – 0.039)      | 48   | 1.5 | 0.14                  |  |
| $b(635)$  | $k(880)$ | 0.36 | 0.043     | 0.016 | (0.011 – 0.076)       | 48   | 2.7 | 0.01                  |  |

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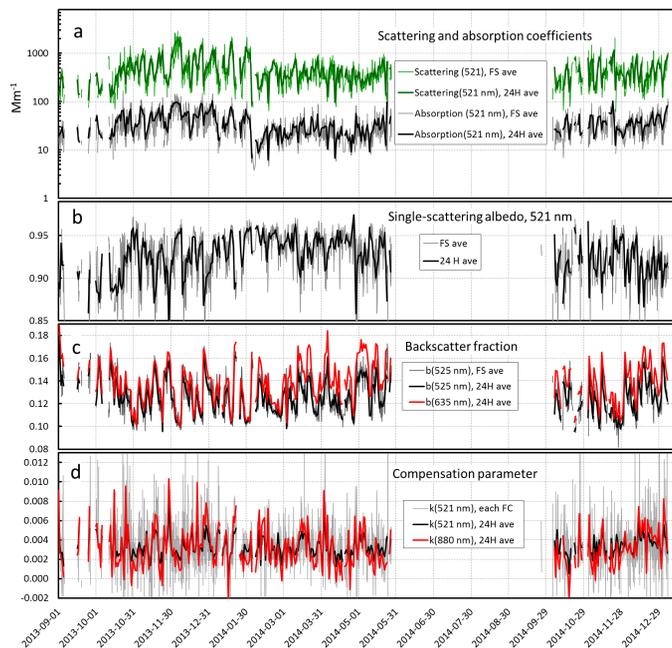


**Table 4.** Regression statistics ( $y = \beta_1 x + \beta_0$ ) of compensation parameter vs. single-scattering albedo at a limited backscatter fraction range. Detailed column description as in Table 1.

| Calculated by using those filter changes during which $b(525 \text{ nm}) = 0.130 \pm 0.005$ |          |      |           |       |                       |      |      |      |  |
|---|----------|------|-----------|-------|-----------------------|------|------|------|--|
| $x$   | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$  | $p$  |  |
| $\omega_0(470)$   | $k(470)$ | 0.08 | 0.009     | 0.006 | (−0.002 – 0.021)      | 411  | 1.7  | 0.10 |  |
| $\omega_0(520)$   | $k(520)$ | 0.08 | 0.011     | 0.006 | (−0.002 – 0.023)      | 411  | 1.6  | 0.10 |  |
| $\omega_0(660)$   | $k(660)$ | 0.09 | 0.014     | 0.008 | (−0.001 – 0.029)      | 411  | 1.9  | 0.06 |  |
| $\omega_0(880)$   | $k(880)$ | 0.13 | 0.024     | 0.009 | (0.006 – 0.043)       | 411  | 2.6  | 0.01 |  |
| Calculated by using those daily averages during which $b(525 \text{ nm}) = 0.130 \pm 0.005$ |          |      |           |       |                       |      |      |      |  |
| $x$   | $y$      | $r$  | $\beta_1$ | s.e.  | 95 % confidence range | d.f. | $t$  | $p$  |  |
| $\omega_0(470)$   | $k(470)$ | 0.04 | −0.002    | 0.008 | (−0.018 – 0.013)      | 66   | −0.3 | 0.77 |  |
| $\omega_0(520)$   | $k(520)$ | 0.03 | −0.002    | 0.009 | (−0.020 – 0.015)      | 66   | −0.2 | 0.80 |  |
| $\omega_0(660)$   | $k(660)$ | 0.05 | 0.004     | 0.010 | (−0.015 – 0.024)      | 66   | 0.4  | 0.68 |  |
| $\omega_0(880)$   | $k(880)$ | 0.18 | 0.017     | 0.012 | (−0.006 – 0.041)      | 66   | 1.5  | 0.14 |  |

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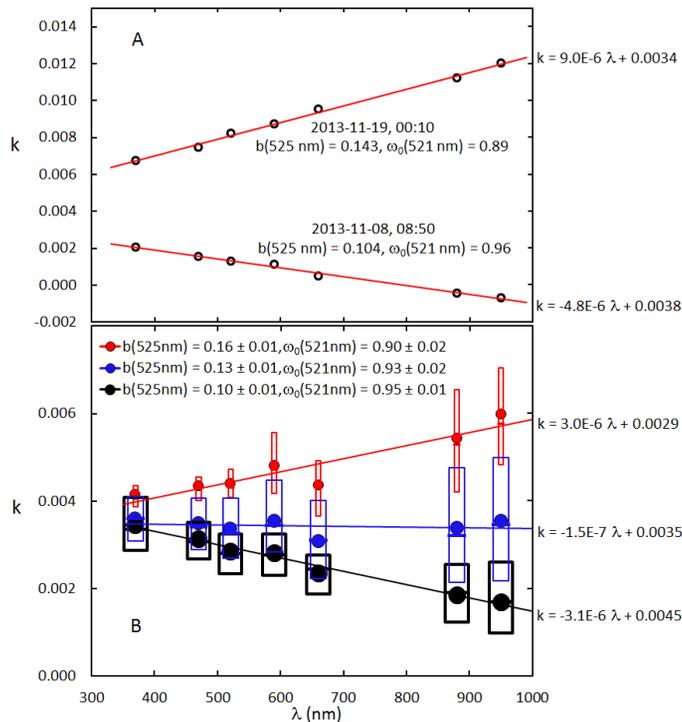


**Figure 1.** Overview of the data. **(a)** Scattering and absorption coefficients at  $\lambda = 520$  nm, **(b)** single-scattering albedo at  $\lambda = 520$  nm, **(c)** backscatter fraction at  $\lambda = 525$  nm and  $\lambda = 635$  nm, and **(d)** the compensation parameter ( $k$ ) at  $\lambda = 520$  nm and at  $\lambda = 880$  nm. In **(a–c)** the thin lines show the filter-spot-averaged (FS ave) values and in **(d)** the individual compensation parameters at each filter change (FC). The thick lines show the 24 h-averaged values in all figures.

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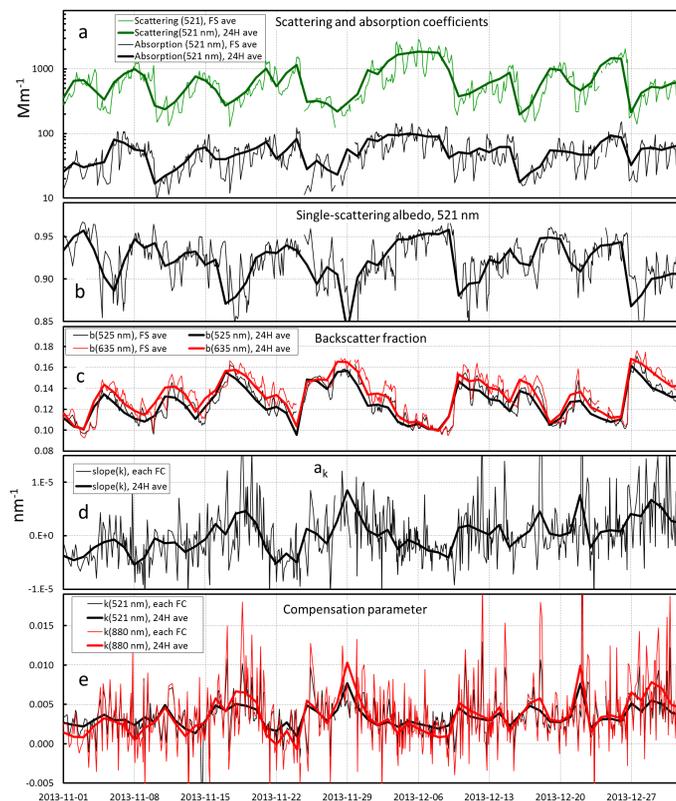


**Figure 2.** Wavelength dependency of the  $k$  factor **(a)** on 08 and 19 November 2013 and **(b)** in the whole data after classification into three bins of green backscatter fraction. In **(b)** the box plots present the 25th to 75th percentiles, the middle lines the medians and the circles the averages in each bin. In **(a)** the lines represent linear regression fittings to the individual  $k$  factors and in **(b)** linear regression fittings to the average  $k$  in each wavelength and backscatter bin.

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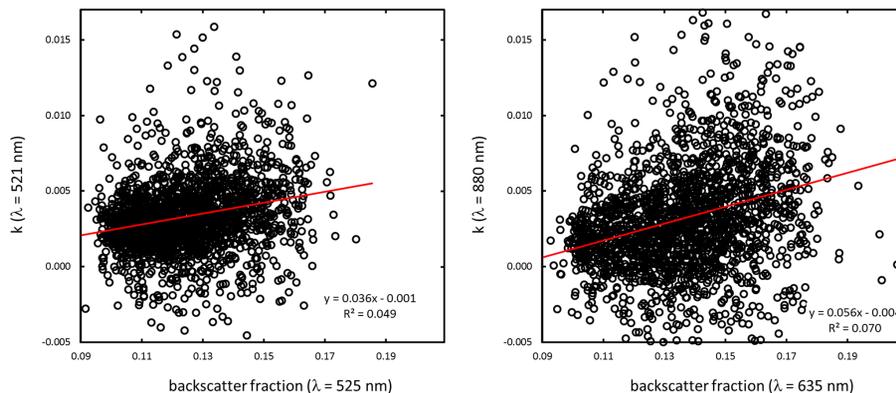
**Figure 3.** Selected optical properties in November–December 2013. **(a)** Scattering and absorption coefficients at  $\lambda = 520$  nm, **(b)** single-scattering albedo at  $\lambda = 520$  nm, **(c)** backscatter fraction at  $\lambda = 525$  nm and  $\lambda = 635$  nm, **(d)** the slope ( $a_k$ ) of the wavelength dependency of the compensation parameter, and **(e)** the compensation parameter ( $k$ ) at  $\lambda = 520$  nm and  $\lambda = 880$  nm. In **(a–c)** the thin lines show the filter-spot-averaged (FS ave) values, in **(d–e)** the individual slopes and compensation parameters at each filter change (FC). The thick lines show the 24 h-averaged values in all figures.

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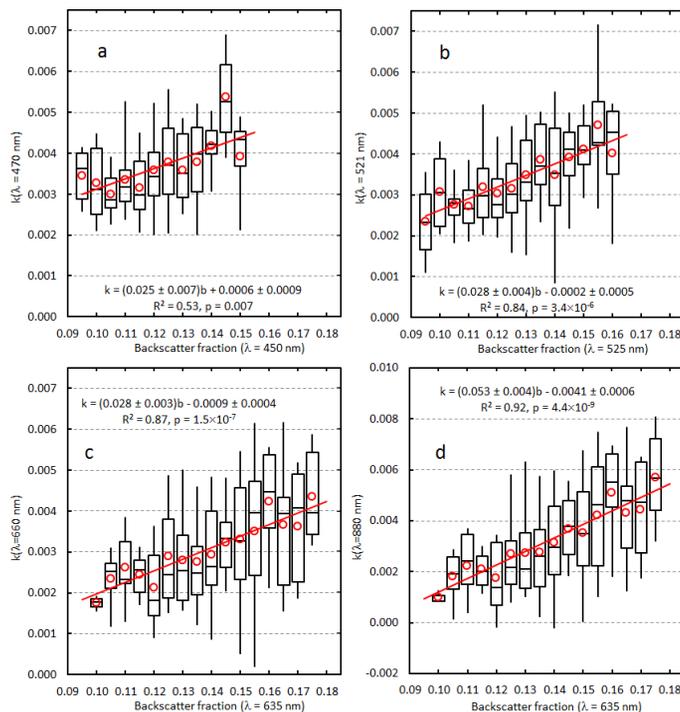


**Figure 5.** The compensation parameter ( $k$ ) of individual filter spot changes calculated for green ( $\lambda = 520$  nm) and near-infrared ( $\lambda = 880$  nm) light as a function of filter-spot-averaged backscatter fraction at the nearest nephelometer wavelengths  $\lambda = 525$  nm and  $\lambda = 635$  nm.

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**Figure 6.** Daily-averaged compensation parameters of blue ( $\lambda = 470 \text{ nm}$ ), green ( $\lambda = 520 \text{ nm}$ ), red ( $\lambda = 660 \text{ nm}$ ), and near-infrared ( $\lambda = 880 \text{ nm}$ ) light classified into 0.005-wide bins of backscatter fraction at the nearest nephelometer wavelengths ( $\lambda = 450, 525$ , and  $635 \text{ nm}$ ). The box plots present the 5th, 25th, 50th, 75th, and 95th percentiles and the circles the averages in each bin. The lines are linear fittings to the bin averages and the uncertainties of the slope and offset the standard errors obtained from the fitting.

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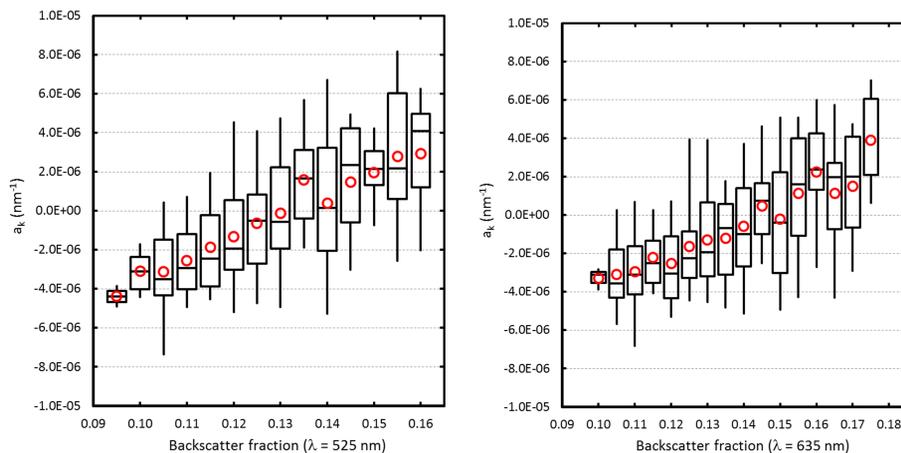
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**Figure 7.** Daily-averaged slope ( $a_k$ ) of the compensation parameter classified into 0.005-wide bins of backscatter fraction of green and red light ( $\lambda = 525 \text{ nm}$ , and  $635 \text{ nm}$ ). The box plots present the 5th, 25th, 50th, 75th, and 95th percentiles and the circles the averages in each bin.

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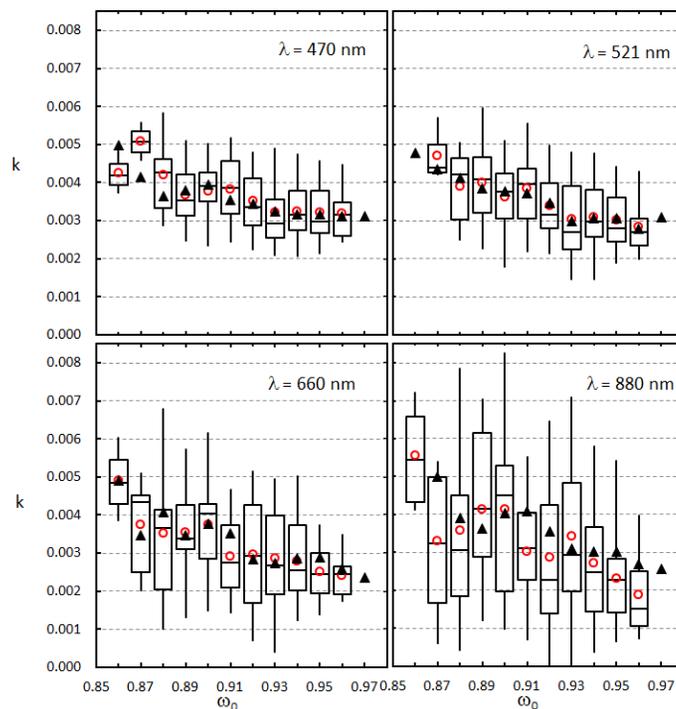
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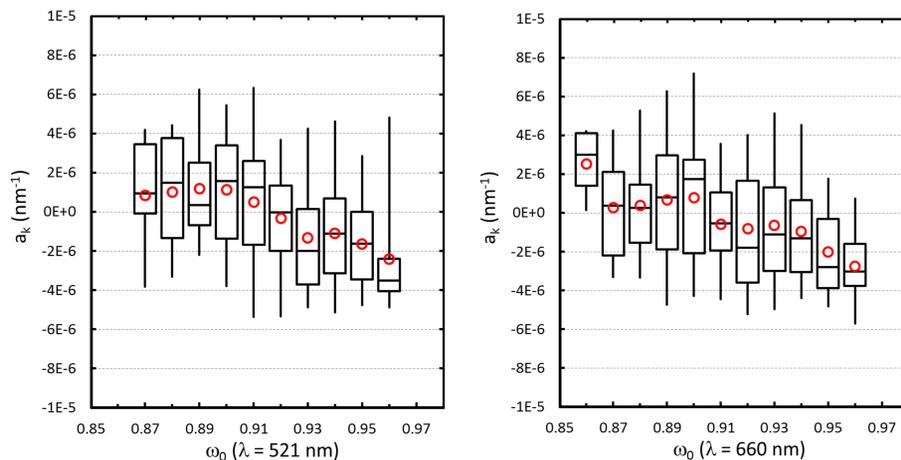


**Figure 8.** Daily-averaged compensation parameters of blue ( $\lambda = 470$  nm), green ( $\lambda = 520$  nm), red ( $\lambda = 660$  nm), and near-infrared ( $\lambda = 880$  nm) light classified into 0.01-wide bins of single-scattering albedo ( $\omega_0$ ) at the same wavelengths. The box plots present the 5th, 25th, 50th, 75th, and 95th percentiles and the circles the averages in each bin. Most  $\omega_0$  bins are based on  $\sigma_{ap}$  calculated with the Collaud Coen et al. (2010) algorithm, the black triangles are the averages of compensation parameters classified into  $\omega_0$  bins with the  $\sigma_{ap}$  calculated by using the Arnott et al. (2005) algorithm.

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**Figure 9.** Daily-averaged slope ( $a_k$ ) of the compensation parameter classified into 0.01-wide bins of single-scattering albedo ( $\omega_0$ ) at green and red wavelengths light ( $\lambda = 520$  and  $660$  nm). The box plots present the 5th, 25th, 50th, 75th, and 95th percentiles and the circles the averages in each bin.

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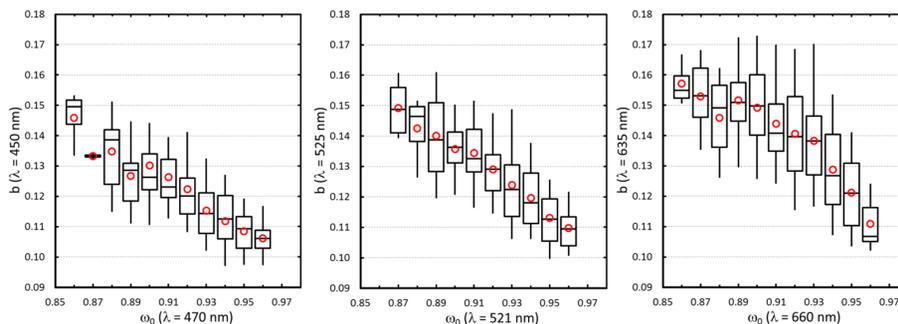
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**Figure 10.** Daily-averaged backscatter fraction ( $b$ ) of blue, green and red light classified into 0.01-wide bins of single-scattering albedo ( $\omega_0$ ). Note:  $b$  is that measured at the nephelometer wavelengths 450, 525 and 635 nm and  $\omega_0$  is that at the aethalometer wavelengths 470, 520 and 660 nm. The box plots present the 5th, 25th, 50th, 75th, and 95th percentiles and the circles the averages in each bin.

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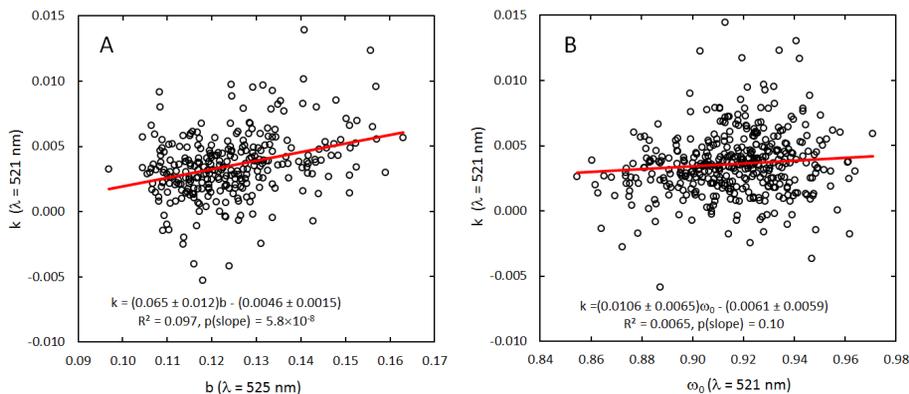
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**Figure 11.** Compensation parameters at the aethalometer green wavelength ( $\lambda = 520$  nm) as a function of **(a)** backscatter fraction and **(b)** single-scattering albedo  $\omega_0$  when  $b$  and  $\omega_0$  were in a narrow range: **(a)** contains data from those filter changes during which the filter-spot-averaged  $\omega_0$  (520 nm) =  $0.930 \pm 0.005$ . **(b)** contains data from those filter changes during which the filter-spot-averaged  $b$ (525 nm) =  $0.130 \pm 0.005$ . The red lines and the equations represent linear regressions fitted to the data.

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