

1 **The Wuhan Atmospheric Radio Exploration**  
2 **(WARE) Radar: Implementation and Initial Results**

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9

10 **Abstract**

11 The recently constructed Wuhan Atmospheric Radio Exploration (WARE) radar is the  
12 first mesosphere-stratosphere-troposphere (MST) radar in the mainland of China,  
13 located at Chongyang, Hubei Province (114°8'8"E, 29°31'58"N, ~ 23° geomagnetic  
14 latitude). WARE radar has a capability of probing the structure and dynamics of the  
15 atmosphere at the altitudes from 3 km to 100 km (excluding 25 km – 60 km). With  
16 fine temporal and spatial resolution, WARE radar provides an outstanding  
17 opportunity ~~for the first time~~ to extensively and intensively investigate various  
18 atmospheric phenomena at the regions of mid-latitude China. In this paper, we present  
19 the main configuration and technical specifications of WARE radar system. For the  
20 first time, we report some initial results: (1) wind field observation from 69 km to 85  
21 km, and wind field observation from 3.2 km to 16.9 km with comparison with results

22 of rawinsonde; (2) tropopause height determined by radar echo power and comparison  
23 between radar tropopause and rawinsonde tropopause (3) atmospheric gravity waves  
24 in the troposphere with ~~—wave length~~wavelength and propagation direction analyzed  
25 by hodograph method; (4) aspect sensitivity of echo power at six specified heights in  
26 the troposphere and stratosphere; (5) diurnal and semi-diurnal tides in tropospheric  
27 and low stratospheric height analyzed by Lomb-Scargle periodogram method.

28

## 29 **1 Introduction**

30 In the past 30 years, mesosphere-stratosphere-troposphere (MST) radars have been  
31 developed and installed around the world. MST radar observes the echoes  
32 backscattered from refractive index fluctuations in the neutral atmosphere and  
33 ionosphere, which can be applied to operational routine observation and scientific  
34 research. The first Very-high-frequency (VHF) MST radar in Jicamaca successfully  
35 observed the atmospheric echoes in 1970s (Woodman and Guillen, 1974). Since then,  
36 studies and constructions of MST radars have shown great growth. For example, the  
37 SOUSY radars (Czechowsky et al., 1984; Rüster et al., 1986), the Esrange radar  
38 (Clison et al., 1999) and the recent MARRSY MST radar (Latteck et al., 2012) in  
39 Norway have shown great capacities for atmospheric research in Europe. VHF MST  
40 radars have also developed widely in the United States since 1970s (Gage and Balsley,  
41 1978; Gage and Green, 1978; Hocking et al., 2001). In Asia, there have been several  
42 VHF-MST radars, such as the middle and upper atmosphere (MU) radar (Fukao et al.,

43 1980, 1990), the Chung\_Li radar (Röttger et al., 1990), the Gadanki radar (Rao et al.,  
44 1995; ~~Jain~~Jain et al., 1995), the Equatorial Atmosphere Radar (EAR) radar (Fukao et  
45 al., 2003), and so on, which have provided numerous important results and findings to  
46 further our understandings of the atmosphere and the ionosphere.

47 Enormous progresses have been made into the atmospheric research by VHF-MST  
48 radars (see the reviews by Hocking (1997; 2001), Fukao (2007) and the references  
49 therein). One of the most important purposes of MST radars is routinely continuous  
50 monitoring the three-dimensional atmospheric winds at altitudes above the boundary  
51 layer. Doppler Beam Swinging (DBS) and Spaced Antenna (SA) methods are the two  
52 primary methods for deriving wind field estimation of MST radars. In particular, MST  
53 radars is capable of measuring the profiles of vertical wind velocities. Conventionally,  
54 the tropopause height is determined from atmospheric temperature by rawinsonde  
55 observations. MST radars provide an effective and efficient approach to estimate the  
56 detailed temporal variation of tropopause, which is very important for studies of  
57 atmospheric dynamics. Due to the fine temporal resolution of three dimensional wind  
58 measurement at different height, MST radars are the most suitable instruments for  
59 atmospheric gravity waves (AGWs) studies, including intrinsic frequencies,  
60 wavelengths, wavenumber spectra, and possible wave sources. Studies of aspect  
61 sensitivity of echo power have always been a significant topic for MST radar  
62 community. For MST research, aspect sensitivity means angle dependence of echo  
63 power, which can be employed to investigate the primary mechanism (scattering or

64 reflection) of the MST radar echoes. With the fine temporal resolution, MST radar is  
65 also suitable for tidal analysis and nonlinear coupling between different atmospheric  
66 waves and oscillations such as AGWs, planetary waves (PW), and quasi biennial  
67 oscillation (QBO).

68 Although considerable development of technology and applications for MST  
69 studies have been made, new instruments and facilities are still planned and  
70 proceeded for further atmospheric research. Recently, Wuhan Atmospheric Radio  
71 Exploration (WARE; 114°8'8"E, 29°31'58"N) radar has been successfully developed  
72 and put into operation (Zhao et al., 2013). The WARE radar is a significant facility of  
73 the Meridian Space Weather Monitoring Project of China (Wang, 2010) that conduct a  
74 comprehensive multi-layered and inter-disciplinary survey and exploration of space  
75 environment by advanced ground based techniques. The WARE radar has great  
76 potential to yield new findings, especially for regional atmospheric characteristics,  
77 due to the capabilities for comprehensive atmospheric research.

78 In this report the design and the implementation of WARE radar system is  
79 summarized. For the first time we present a number of initial observations of WARE  
80 radar, including measurements of tropospheric and mesospheric wind, determination  
81 of the tropopause, studies of AGWs, tropospheric aspect sensitivity and atmospheric  
82 tides investigation.

83

## 84 **2 Brief Description of WARE Radar System**

85 The WARE radar is the first MST radar put into service in mainland of China,  
86 which is located at Chongyang, Hubei Province of China (114°8'8"E, 29°31'58"N)  
87 with the geomagnetic latitude of  $\sim 23^\circ$ . The altitude of the radar site is 62 meters  
88 above the sea level. The inclination and declination angles of geomagnetic field are  
89 approximately  $44.7^\circ$  and  $-3.7^\circ$  respectively at the altitude of 110 km above the radar  
90 site.

91 The WARE radar is a fully distributed, all solid state, and coherent pulsed  
92 Doppler radar operating at 53.8 MHz with an average power aperture product of  $2.3 \times$   
93  $10^8 \text{ Wm}^2$ . The reliable detection height range covers from roughly 3 km to 100 km  
94 (not including 25 km – 60 km). The phased array of WARE radar consists of  $24 \times 24$   
95 three-element Yagi-Uda antennas evenly distributed over a total area of 10000 m<sup>2</sup>.  
96 The radar beam can be steered into five directions (East, West, South, North, and  
97 vertical) independently, changing continuously from vertical to  $20^\circ$  off-zenith angle  
98 with a step of  $1^\circ$  from pulse to pulse. The echoes from five beams are collected  
99 alternately by the receiver. DBS and active phased array technique are utilized to  
100 probe the structure and the dynamics of the atmosphere. WARE radar is designed to  
101 operate at three independent modes which correspond to monitoring the troposphere  
102 (low mode, 3.5 km – 10 km), the low stratosphere (medium mode 10 km – 35 km),  
103 and the mesosphere (high mode 60 km – 90 km), respectively.

104 Characteristics of atmospheric turbulence echoes, including the Doppler

105 velocity and spectrum width, are estimated after removal of background noise and  
106 clutter. Distinct from time domain pulse radar system, clutter suppression is  
107 accomplished by spectrum analysis due to frequency domain detection of WARE  
108 radar. The technical parameters and performance parameters of three modes are  
109 tabulated in Table 1, respectively. More detailed description about WARE system and  
110 hardware can be found in Zhao et al. (2013).

111

## 112 **3 Preliminary Results of WARE Radar Observations**

### 113 **3.1 Observations of atmospheric wind field**

114 Measurement of fine quality wind field is one of the major objectives of MST  
115 radars (~~Gage and Vanzandt, 1981; Balsley, 1983~~). WARE radar has the capability of  
116 measuring the wind field based upon the DBS technique. Comparisons of wind field  
117 observations between WARE radar and GPS rawinsonde were carried out  
118 immediately after the WARE radar was established. Figure 1 (a) presents the power  
119 spectrum of radar echoes received by the east, west, south, and north beams tilted to  
120 20° off-zenith angle, and Figure 1 (b) shows the comparison of radar wind and  
121 rawinsonde wind obtained at 16:00 LT on September 11, 2011. Clearly, the  
122 observations of WARE radar are reasonably consistent with the rawinsonde  
123 observations for both wind velocity and wind direction, indicating that our newly  
124 constructed WARE radar works properly and efficiently to provide reliable  
125 observations for further studies. One of the potential capability of MST radar is to

126 provide the direct wind measurement of vertical velocity. Figure 2 presents vertical  
127 velocity for continuous 24 hours on October 2, 2012.

128 An example of mesospheric wind field measurements from 69 to 85 km obtained  
129 by WARE radar on March 15, 2011 is shown in Figure 23. Figure 2-3(a) presents the  
130 power spectrum received by east, west, south, and north beams tilted to 20 °off-zenith  
131 angle. Up to 20 ° off-zenith beams are utilized for the mesospheric wind estimation,  
132 which cover a vast horizontal area and get enough horizontal wind velocity  
133 component. However, these measurement need more homogenous backscatter in the  
134 observation volume (Browning and Wexler 1967; Waldteufel and Corbin 1973; Stober  
135 et al., 2013).

136 Figure 2-3(b) presents the meridional and zonal wind estimated by echo power  
137 spectrum. Lee et al. (2014) suggest an influence from local mountain waves. The  
138 WARE radar is located at the Jiangnan Plain, no large mountain ridges are close to the  
139 radar site. Small uncontinuous foothills are located about 50 km south of the radar,  
140 which could produce mountain lee waves. These potential waves could induce  
141 uncertainties in radar wind measurement. (Stober et al., 2011)

142 –It is expected that the long-term observation by WARE radar will establish a  
143 unique database to investigate the profiles of mesospheric wind in the mid-latitude  
144 region of China.

### 146 **3.2 Observations of the tropopause**

147 Tropopause is a natural stable layer which plays a significant role in the  
148 stratosphere-troposphere exchange (STE). Characteristics of tropopause have been  
149 studied in many ways and in many regions (Reid and Gage, 1996; Hermawan et al.,  
150 1998; Yamamoto et al., 2003; Das et al., 2008; Mehta et al., 2008, 2011). The detailed  
151 time variation of tropopause structure is very important for studies of dynamical  
152 atmospheric properties. WARE radar provides a good opportunity to study the  
153 mid-latitude tropopause.

154 Conventionally, the height of tropopause can be determined by the lapse rate (lapse  
155 rate tropopause, LRT) (World Meteorological Organization (WMO), 1996), cold point  
156 (cold point tropopause, CPT) (Selkirk, 1993), and radar echo power (radar tropopause,  
157 RT) (Gage and Green, 1979, 1982; Hall et al., 2009). In our present study, we simply  
158 adopt refractivity structure constant ( $C_n^2$ ) derived from radar echo power to evaluate  
159 the tropopause location (Rao et al., 1997; Ghosh et al., 2001; Zink et al., 2004). Two  
160 experiments are shown in Figure 3(a), which were executed around 17:15 LT on  
161 September 10, 2011 and around 07:21 LT on September 11, 2011 for validation of the  
162 radar measurements. The green line in Figure 3(a) is the  $C_n^2$  value estimated from  
163 radar echo power. The location of maximum  $C_n^2$  value implies the height of  
164 tropopause, which was about 16 km in the two cases. The blue points are the recorded  
165 temperature derived from rawinsonde and the blue line is the fitted temperature  
166 profile. The coldest point of the temperature is the height of tropopause determined by  
167 rawinsonde, which was about 17 km. The location of tropopause can be different

168 according to different tropopause definitions and tropopause dynamics. (Yamamoto et  
169 al., 2003; Das et al., 2008) In addition, the tropopause is not just a thin layer but a  
170 transition region between the troposphere and stratosphere (Mehta et al., 2008).  
171 Figure 4 (a) demonstrate the difference between the cold point tropopause (CPT) from  
172 rawinsonde and radar tropopause. It should also be noted that the height resolution is  
173 0.6 km in the medium operational mode, which add some uncertainties for radar  
174 tropopause location. Considering that the height resolution of WARE radar operated at  
175 medium mode is 0.6 km, the heights of tropopause derived from radar and rawinsonde  
176 are consistent with each other basically.

177 After validation of radar tropopause, a 20.5-hour observation was carried out from  
178 08:05 LT, January 2, 2012 to 04:35 LT, January 3, 2012. Figure ~~3-4~~(b) shows the  
179 result. As seen, the height of radar tropopause was between 11 km to 12 km during  
180 this time period, which was much lower than that 1 shown in Fig. ~~3-4~~(a). Variation in  
181 tropopause height can be attributed to different seasons for the radar observations. The  
182 tropopause height is usually higher in summertime than in wintertime.

183

### 184 **3.3 Observations of atmospheric gravity waves**

185 Atmospheric gravity waves (AGWs) play an important role in transporting energy  
186 and momentum, in contributing turbulence and mixing, and in influencing the mean  
187 circulation and thermal structure of the middle atmosphere (Fritts and Alexander,  
188 2003). In addition, AGWs are crucial to understand the coupling process of the upper

189 atmosphere and the ionosphere (Hines, 1960). Quasi-monochromatic (QM) AGWs are  
190 frequently observed with airglow imagers, lidars, radars and rawinsondes. On this  
191 aspect, WARE radar provides a standard and effective tool to study these AGWs.  
192 Hodograph method (Gavrilov et al., 1996; 1997; Hu et al., 2002; Zhang and Yi, 2005)  
193 is utilized here to extract the parameters of dominant QM AGWs. To obtain the  
194 parameters of AGW, the background winds are firstly removed by fitting a second  
195 order polynomial to the horizontal wind profiles (Zhang and Yi, 2005). The results  
196 shown in Figure 4 are two typical examples of QM AGWs observed by WARE radar  
197 at 06:05 LT and 13:05 LT, on September 26, 2011. Figure 4-5(a) and (b) show wave  
198 fits of vertical profiles of meridional wind disturbance at heights from 3.04 to 9.85 km  
199 at two respective scenarios, while Figure 4-5(c) and (d) show the zonal component.  
200 Figure 4-5(e) and (f) present the vertical wind disturbance. The ~~wavelength~~  
201 ~~lengths~~ of the two QM AGWs are calculated to be 4.99 km and 3.25 km, respectively.  
202 Hodographs of the meridional wind versus zonal wind are illustrated in Figure 4-5(g)  
203 and (h). The rotations are both anti-clockwise, indicating downward propagating  
204 waves.

205 By analyzing the polarization relation of AGW, the AGW horizontal propagation  
206 directions are parallel to the major axis of the ellipse in hodographs. We can estimate  
207 the propagation directions are  $23^\circ$  or  $23^\circ+180^\circ$  and  $45^\circ$  or  $45^\circ+180^\circ$ , since the  
208 directions have  $180^\circ$  ambiguity. In order to resolve this ambiguity, the simultaneous  
209 measurement of the temperature profile can be utilized by radiosonde observations or

210 Na lidar (Hu et al., 2002; Zhang and Yi, 2005), However WARE radar can not provide  
211 the temperature profile directly, the ambiguity can also be eliminated by the vertical  
212 wind profile (Muraoka et al., 1987; Tsuda et al., 1990). Hence the horizontal  
213 propagation direction can be determined by combination of the hodograph method  
214 and the measurement of vertical velocity of wind. For the cases presented in Figure 4,  
215 the horizontal propagation directions are finally identified as 23 °(left column) and  
216 225 °(right column), which are clockwise to the north. The wave intrinsic frequencies  
217 of these QM AGWs are also calculated according to the hodographs, which is 10.3  
218 hour and 7.2 hour, the local Coriolis frequency is  $7.29 \times 10^{-5} \text{ rads}^{-1}$ . These QM AGWs  
219 in the upper troposphere have been frequently observed by WARE radar, a statistical  
220 analysis of which is our undertaking work.

221

### 222 **3.4 Observations of aspect sensitivity of echo power**

223 Aspect sensitivity means the dependence of echo power on antenna beam pointing  
224 angle, which has been reported by several researchers (Gage and Green, 1978; Röttger  
225 and Liu, 1978; Tsuda et al., 1986). MST radar echoes are known to be aspect sensitive  
226 due to anisotropic backscattering and Fresnel reflection/scattering, which in turn  
227 influence the determination of wind components and measurement of turbulence  
228 parameters. Jain et al. (1997) suggested that for smaller beam angles the horizontal  
229 wind component may be underestimated by as much as 30%. Therefore several  
230 theoretical models are proposed to explain scattering/reflection mechanisms in order

231 to facilitate the understanding of aspect sensitivity and its effect on the radar echoes.

232 WARE radar operates at five beams including one vertical beam and four oblique  
233 beams. Each oblique beam has been designed to be able to scan from vertical to 20°  
234 continuously with a step of 1°. Therefore WARE radar has the capability for  
235 investigating the zenith angle dependence of the backscattered echo power at different  
236 altitudes. Figure 5-6 shows the variation of echo power as a function of beam zenith  
237 angle at six specified heights from 4.75 km to 17.85 km. Observations are carried out  
238 at 13:13 LT, December 27, 2011.

239 Figure 5-6 shows that (1) aspect sensitivity occurs throughout the troposphere and  
240 low stratosphere; (2) aspect sensitivity is higher in the troposphere than in the  
241 stratosphere in this observation; (3) the relative echo power for smaller zenith angles  
242 ( $\leq 10^\circ$ ) decreases faster than the larger angles ( $> 10^\circ$ ). Usually, the degree of aspect  
243 sensitivity is lower in the troposphere than in the stratosphere. This is due to turbulent  
244 air in the troposphere and stable atmosphere in the stratosphere. There are several  
245 interpretations on the causative mechanisms of aspect sensitivity. Fresnel  
246 reflection/scattering and anisotropic scattering are the two leading suggestions.  
247 (Röttger and Liu, 1978; Crane, 1980; Doviak and Zrnic, 1984; Woodman and Chu,  
248 1989) Study of aspect sensitivity can yield a deeper understanding of the mechanisms  
249 of the atmospheric radar echoes, like isotropic/anisotropic turbulence or Fresnel  
250 reflecting/scattering structures. However, it is difficult to separate the respective  
251 contributions of these echoing mechanisms. In this case, we present irregular pattern

252 of aspect sensitivity at six specified heights of troposphere and low stratosphere,  
253 which could be an evidence for the complicated interpretation of atmospheric aspect  
254 sensitivity and that combined mechanism could make contributions. Radar echoes  
255 also could be come from different scatters, especially at low beam angles (Chen and  
256 Furumoto, 2013). Tsuda et al. (1997) suggest that horizontal component of  
257 atmospheric gravity wave motions could distort the refractivity surface. This  
258 mechanism could also generate irregular pattern of aspect sensitivity.~~Our observation~~  
259 ~~indicate that the mechanism for aspect sensitivity could be complicated and Different~~  
260 ~~interpretations should be taken into consideration.~~

261

### 262 **3.5 Observations of diurnal and semi-diurnal tides**

263 Atmospheric tides are defined as atmospheric waves or oscillations with periods of  
264 harmonics of a solar day. The subset migrating tides propagate westward with zonal  
265 wavenumbers equal to the frequencies in cycles per day. The migrating tides are  
266 generated primarily by ozone heating in the stratosphere and water vapor heating in  
267 the troposphere, which are uniform zonal distributed. The zonal wavenumbers of  
268 nonmigrating tides are not equal to the frequencies in cycles per day, which have  
269 more local characteristics. Tides are believed to play an important role in large scale  
270 circulation patterns and the dynamics of mesosphere and lower thermosphere. (Forbes,  
271 1982; Forbes et al., 1997; Vincent et al., 1998; Manson et al., 2002) However, due to  
272 the small amplitudes of tides in the troposphere and lower stratosphere, the studies on

273 tides in this region require more efforts and deserve further investigation.

274 We report diurnal and semi-diurnal tide analysis by WARE radar observations in  
275 the lower atmosphere from October 1 to October 12, 2011. The observations of  
276 WARE radar provide 30 minutes time resolution and allow us to analyze in the  
277 temporal domain. Meridional and zonal wind are firstly derived from power spectrum  
278 of radar echoes. Figure 6-7(a) and Figure 7-8(a) show meridional and zonal wind  
279 from 3 km to 25 km with height resolution of 150 meters during continuous 12 days.  
280 Then the perturbation wind field is obtained by subtracting the background wind. We  
281 calculate the background wind by applying a second-order polynomial fitting to the  
282 vertical profiles of horizontal and vertical winds respectively, following Zhang and Yi  
283 (2005). After that, a high pass filter with a cutoff at 36 h is applied to remove the  
284 influence of planetary waves (PW) and quasi biennial oscillation (QBO). Finally, a  
285 Lomb-Scargle periodogram analysis (Scargle, 1981; 1982) is performed on the  
286 resultant time series data. Significant values in Figure 6-7(b) and Figure 7-8(b) can  
287 be found around the period of 12 hour and 24 hour, which indicate semi-diurnal and  
288 diurnal oscillation. It should also be noted that at the low altitude the values  
289 concentrate on the 12 hour and 24 hour, while at the high altitude the values tend be  
290 scattered around these two periods. The results indicate the influence of secondary  
291 waves generated by the planetary/tidal wave interaction. (Huang et al., 2009)\_

292 Tidal horizontal wind amplitudes can reach to several tens of m/s in mesosphere  
293 and low thermosphere, which have been intensively studied for thirty years. However,

294 due to the weak amplitude, tides in troposphere and lower stratosphere have not been  
295 subjected to sufficient study. Recent research with intense rawinsonde and radar  
296 measurement [Whiteman et al, 1996; Huang et al, 2009; Sakazaki et al, 2010] have  
297 presented tidal wind perturbations in tropospheric and low stratospheric data.  
298 However, due to nonlinear wave-flow and wave-wave interactions, tidal frequency  
299 spectral estimation could be biased with periods around the diurnal and semi-diurnal  
300 tides.

301

#### 302 **4. Summary**

303 Here in this paper we have summarized the design and implementation of WARE  
304 radar – the first VHF-MST radar in the mainland of China. A number of initial results  
305 regarding atmospheric wind field, tropopause, atmospheric gravity waves (AGWs),  
306 aspect sensitivity of echo power, and atmospheric tides have been reported. These  
307 results demonstrate the outstanding capabilities of WARE radar for comprehensive  
308 atmospheric researches. We have a strong feeling that as a unique VHF-MST radar  
309 facility at mid-latitude in the mainland of China and as an integrated part of  
310 well-planned Chinese Meridian Space Weather Monitoring Project, WARE radar has  
311 great potentials to yield new findings especially of the regional atmospheric  
312 characteristics. Further system improvements and expansions will include radio  
313 acoustic sounding system (RASS) to obtain local temperature profiles. Combinations  
314 of AGWs observations by WARE radar with simultaneous GPS network and HF

315 Doppler observations is also able to provide profound clues to track the correlations  
316 between AGWs and TIDs. Therefore, this kind of study can greatly enhance our  
317 knowledge of the neutral atmosphere-ionosphere coupling process, especially at low-  
318 and middle-latitudes. Collaborative experiment campaigns of our WARE MST radar,  
319 for instance, during the deep convection, cold front, Mei-Yu season, and / or magnetic  
320 storms, together with the MU radar and the Chung-Li radar, all of which are located in  
321 East Asia, will establish a truly unique platform for in-depth investigations and  
322 comprehensive understandings of the structures and the dynamics of the Earth's  
323 atmosphere.

324

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332 Data Centre for Meridian Space Weather Monitoring Project (<http://159.226.22.74/>).  
333 The rawinsonde data can be provided by corresponding author via private  
334 communication.

335

336 **Reference**

337 Balsley, B. B., Ecklund, W. L., Carter, D. A., and Johnston, P. E.: The MST radar at  
338 Poker Flat, Alaska, *Radio Sci.*, 15(2), 213–223, doi:10.1029/RS015i002p00213,  
339 1980.

340 ~~Balsley, B. B.: Poker Flat MST radar measurement of winds and wind variability in  
341 the mesosphere, stratosphere, and troposphere, *Radio Sci.*, 18(6), 1011–1020,  
342 doi:10.1029/RS018i006p01011, 1983.~~

343 Browning, K. and Wexler, R.: The Determination of Kinematic Properties of a Wind  
344 field Using Doppler Radar, *J. Appl. Meteorol.*, 7, 105–113, 1968.

345 Chen, Jenn-Shyong and Furumoto, Jun-ichi: Measurement of Atmospheric Aspect  
346 Sensitivity Using Coherent Radar Imaging after Mitigation of Radar Beam  
347 Weighting Effect. *J. Atmos. Oceanic Technol.*, 30, 245–259, 2013.

348 Chilson, P. B., Kirkwood, S., and Nilsson, A.: The Erange MST radar: A brief  
349 introduction and procedure for range validation using balloons, *Radio Sci.*, 34(2),  
350 427–436, doi:10.1029/1998RS900023, 1998.

351 Chu, Y. H., Chao, J. K., Liu, C. H., and Röttger, J.: Aspect sensitivity at tropospheric  
352 heights measured with vertically pointed beam of the Chung-Li VHF radar,  
353 *Radio Sci.*, 25(4), 539–550, doi:10.1029/RS025i004p00539, 1990.

354 Crane, R. K.: A review of radar observations of turbulence in the lower stratosphere,  
355 *Radio Sci.*, 15(2), 177–193, doi:10.1029/RS015i002p00177, 1980.

356 Czechowsky, P., Schmidt, G., and Rüter, R.: The mobile SOUSY Doppler radar:

357 Technical design and first results, *Radio Sci.*, 19(1), 441–450,  
358 doi:10.1029/RS019i001p00441, 1984.

359 Das, S. S., Jain, A. R., Kumar, K. K., and Rao, Narayana D.: Diurnal variability of the  
360 tropical tropopause: Significance of VHF radar measurements, *Radio Sci.*, 43,  
361 RS6003, doi:10.1029/2008RS003824, 2008.

362 Doviak, R. J., and Zrnic', D. S.: Reflection and scatter formula for anisotropically  
363 turbulent air, *Radio Sci.*, 19(1), 325–336, doi:10.1029/RS019i001p00325, 1984.

364 Forbes, J. M.: Atmospheric tides: 1. Model description and results for the solar diurnal  
365 component, *J. Geophys. Res.*, 87(A7), 5222–5240, doi:10.1029/  
366 JA087iA07p05222, 1982.

367 Forbes, J. M., Hagan, M. E., Zhang, X., and Hamilton, K.: Upper atmosphere tidal  
368 oscillations due to latent heat release in the tropical troposphere, *Ann. Geophys.*,  
369 15, 1165–1175, 1997.

370 Fukao, S., Kato, S., Aso, T., Sasada, M., and Makihira, T.: Middle and upper  
371 atmosphere radar (MUR) under design in Japan, *Radio Sci.*, 15(2), 225–231,  
372 doi:10.1029/RS015i002p00225, 1980.

373 Fukao, S., Hashiguchi, H., Yamamoto, M., Tsuda, T., Nakamura, T., Yamamoto, M. K.,  
374 Sato, T., Hagio, M., and Yabugaki, Y.: Equatorial Atmosphere Radar (EAR):  
375 System description and first results, *Radio Sci.*, 38, 1053,  
376 doi:10.1029/2002RS002767, 3, 2003.

377 Fukao, S.: Recent Advances in Atmospheric Radar Study, *J. Meteor. Soc. Japan*, 85B,

378 215-239, 2007.

379 Fukao, S., Sato, T., Tsuda, T., Yamamoto, M., Yamanaka, M. D., and Kato, S.: MU  
380 radar: New capabilities and system calibrations, *Radio Sci.*, 25(4), 477–485,  
381 doi:10.1029/RS025i004p00477, 1990.

382 Fritts, D. C., and Alexander, M. J.: Gravity wave dynamics and effects in the middle  
383 atmosphere, *Rev. Geophys.*, 41(1), 1003, doi:10.1029/2001RG000106, 2003.

384 Gage, K. S., and Balsley, B. B.: Doppler radar probing of the clear atmosphere, *Bull.*  
385 *Am. Meteorol. Soc.*, 59, 1074-1093, 1978.

386 Gage, K. S., and Green, J. L.: Evidence for specular reflection from monostatic VHF  
387 radar observations of the stratosphere, *Radio Sci.*, 13(6), 991–1001,  
388 doi:10.1029/RS013i006p00991, 1978.

389 Gage, K. S., and Green, J. L.: Tropopause detection by partial specular reflection with  
390 very-high-frequency radar, *Science*, 203, 1238–1240,  
391 doi:10.1126/science.203.4386.1238, 1979.

392 ~~Gage, K. S., and Vanzandt, T. E.: Wind Measurement Techniques Available for the~~  
393 ~~Middle Atmosphere Program, *J. Geophys. Res.*, 86(C10), 9591–9598,~~  
394 ~~doi:10.1029/JC086iC10p09591, 1981.~~

395 Gage, K. S., and Green, J. L.: An objective method for the determination of  
396 tropopause height from VHF radar observations, *J. Appl. Meteorol.*, 21,  
397 1150–1154, 1982.

398 Gavrilov, N. M., Fukao, S., Nakamura, T., Tsuda, T., Yamanaka, M. D., and

399 Yamamoto, M.: Statistical analysis of gravity waves observed with the middle  
400 and upper atmosphere radar in the middle atmosphere: 1. Method and general  
401 characteristics, *J. Geophys. Res.*, 101(D23), 29511–29521,  
402 doi:10.1029/96JD01447, 1996.

403 Gavrilov, N. M., Fukao, S., Nakamura, T., and Tsuda, T.: Statistical analysis of gravity  
404 waves observed with the middle and upper atmosphere radar in the middle  
405 atmosphere: 2. Waves propagated in different directions, *J. Geophys. Res.*,  
406 102(D12), 13433–13440, doi:10.1029/97JD00736, 1997.

407 Ghosh, A. K., V. Siva Kumar, K. Kishore Kumar, and A. R. Jain (2001), VHF radar  
408 observation of atmospheric winds, associated shears and  $C_2n$  at a tropical  
409 location: interdependence and seasonal pattern, *Ann. Geophys.*, 19, 965-973,  
410 doi:10.5194/angeo-19-965-2001, 2001.

411 Hall, C. M., Röttger, J., Kuyeng, K., Sigernes, F., Claes, S., and Chau, J.: First results  
412 of the refurbished SOUSY radar: Tropopause altitude climatology at 78°N, 16°E,  
413 2008, *Radio Sci.*, 44, RS5008, doi:10.1029/2009RS004144, 2009.

414 Hines, C. O.: Internal atmospheric gravity waves at ionospheric heights, *Can. J. Phys.*,  
415 38, 1441-1481, 1960.

416 Hocking, W. K., Ruster, R., and Chechowsky, P.: Absolute reflectivities and aspect  
417 sensitivities of VHF radio wave scatterers measured with the SOUSY radar, *J.*  
418 *Atmos. Terr. Phys.*, 48, 131–144, 1986.

419 Hocking, W. K.: Recent advances in radar instrumentation and techniques for studies

420 of the mesosphere, stratosphere, and troposphere, *Radio Sci.*, 32(6), 2241–2270,  
421 1997.

422 Hocking, W. K., Kelley, M., Rogers, R., Brown, W. O. J., Moorcroft, D., and St.  
423 Maurice, J.-P.: Resolute Bay VHF radar: A multipurpose tool for studies of  
424 tropospheric motions, middle atmosphere dynamics, meteor physics, and  
425 ionospheric physics, *Radio Sci.*, 36(6), 1839–1857, doi:10.1029/2000RS001005,  
426 2001.

427 Hocking, W. K.: A review of Mesosphere-Stratosphere-Troposphere (MST) radar  
428 developments and studies, circa 1997-2008, *J. Atmos. Solar-Terr. Phys.*, 73(9),  
429 848-882, doi:10.1016/j.jastp.2010.12.009, 2011.

430 Hu, X., Liu, A. Z., Gardner, C. S., and Swenson, G. R.: Characteristics of  
431 quasi-monochromatic gravity waves observed with Na lidar in the mesopause  
432 region at Starfire Optical Range, NM, *Geophys. Res. Lett.*, 29(24), 2169,  
433 doi:10.1029/2002GL014975, 2002.

434 Huang, C. M., Zhang, S. D., and Yi, F.: Intensive radiosonde observations of the  
435 diurnal tide and planetary waves in the lower atmosphere over Yichang (111 °18'  
436 E, 30 °42' N), China, *Ann. Geophys.*, 27, 1079-1095,  
437 doi:10.5194/angeo-27-1079-2009, 2009.

438 Jain, A. R., Rao, Y. Jaya, and Rao, P. B.: Aspect sensitivity of the received radar  
439 backscatter at VHF: Preliminary observations using the Indian MST radar, *Radio*  
440 *Sci.*, 32(3), 1249–1260, doi:10.1029/97RS00252, 1997.

441 Jain, A. R., Rao, Y. J., Rao, P. B., Anandan, V. K., Damle, S. H., Balamuralidhar, P.,  
442 Kulakarni, A., and Viswanathan, G.: Indian MST radar 2. First scientific results  
443 in ST mode, *Radio Sci.*, 30(4), 1139–1158, doi:10.1029/95RS00652, 1995.

444 Latteck, R., Singer, W., Rapp, M., Vandeppeer, B., Renkwitz, T., Zecha, M., and Stober,  
445 G.: MAARSY: The new MST radar on Andøya—System description and first  
446 results, *Radio Sci.*, 47, RS1006, doi:10.1029/2011RS004775, 2012.

447 Manson, A. H., Luo, Y., and Meek, C.: Global distributions of diurnal and  
448 semi-diurnal tides: observations from HRDI-UARS of the MLT region, *Ann.*  
449 *Geophys.*, 20, 1877–1890, 2002.

450 Mehta, S. K., Murthy, B. V. K., Rao, D. N., Ratnam, M. V., Parameswaran, K., Rajeev,  
451 K., Raju, C. S., and Rao, K. G.: Identification of tropical convective tropopause  
452 and its association with cold point tropopause, *J. Geophys. Res.*, 113, D00B04,  
453 doi:10.1029/2007JD009625, 2008.

454 Mehta, S. K., Ratnama, M. V., and Murthyb, B. V. K.: Characteristics of the tropical  
455 tropopause over different longitudes, *J. Atmos. Solar-Terr. Phys.*, 73(17-18),  
456 2462-2473, doi:10.1016/j.jastp.2011.09.006, 2011.

457 Muraoka Y., Kawahira, Sato, K. T., Tsuda, T., Fukao, S., and Kato, S.: Characteristics  
458 of mesospheric internal gravity waves observed by MU Radar. *Geophys. Res.*  
459 *Lett.*, 14(11), 1154–1157, doi: 10.1029/GL014i011p01154, 1987.

460 Rao, P. B., Jain, A. R., Kishore, P., Balamuralidhar, P., Damle, S. H., and Viswanathan,  
461 G.: Indian MST radar 1. System description and sample vector wind

462 measurements in ST mode, *Radio Sci.*, 30(4), 1125–1138,  
463 doi:10.1029/95RS00787, 1995.

464 Rao, D. N., Kishore, P., Rao, T. N., Rao, S. V. B., Reddy, K. K., Yarraiah, M., and  
465 Hareesh, M.: Studies on refractivity structure constant, eddy dissipation rate, and  
466 momentum flux at a tropical latitude, *Radio Sci.*, 32(4), 1375–1389,  
467 doi:10.1029/97RS00251, 1997.

468 Reid, G. C., and Gage, K. S.: The tropical tropopause over the western Pacific: Wave  
469 driving, convection, and the annual cycle, *J. Geophys. Res.*, 101(D16),  
470 21,233–21,241, doi:10.1029/96JD01622, 1996.

471 Röttger, J., and Liu, C. H.: Partial reflection and scattering of VHF radar signals from  
472 the clear atmosphere. *Geophys. Res. Lett.*, 5: 357–360. doi:  
473 10.1029/GL005i005p00357, 1978

474 Röttger, J., et al.: The Chung-Li VHF radar: Technical layout and a summary of initial  
475 results, *Radio Sci.*, 25(4), 487–502, doi:10.1029/RS025i004p00487, 1990.

476 Rüster, R., Klostermeyer, J. and Röttger, J.: SOUSY VHF Radar Measurements in the  
477 Lower and Middle Atmosphere, *IEEE Transactions on Geoscience and Remote*  
478 *Sensing*, GE-24, 6, 966-974, 1986.

479 Scargle, J. D.: Studies in astronomical time series analysis, I – Modeling random  
480 process in the time domain, *Astrophysical J. Supplement Series*, 45, 1–71, 1981.

481 Scargle, J. D.: Studies in astronomical time series analysis, II – Statistical aspects of  
482 spectral analysis of unevenly spaced data, *Astrophysical J.*, 263, 835–853, 1982.

483 Selkirk, H. B.: The tropopause cold trap in the Australian monsoon during  
484 STEP/AMEX 1987. *J. Geophys. Res.* 98, 8591–8610, 1993.

485 [Stober, G., Latteck, R., Rapp, M., Singer, W., and Zecha, M.: MAARSY – the new](#)  
486 [MST radar on Andøya: first results of spaced antenna and Doppler measurements](#)  
487 [of atmospheric winds in the troposphere and mesosphere using a partial array,](#)  
488 [Adv. Radio Sci., 10, 291-298, doi:10.5194/ars-10-291-2012, 2012.](#)

489 [Stober, G., Sommer, S., Rapp, M., and Latteck, R.: Investigation of gravity waves](#)  
490 [using horizontally resolved radial velocity measurements, Atmos. Meas. Tech., 6,](#)  
491 [2893-2905, doi:10.5194/amt-6-2893-2013, 2013.](#)

492 Tsuda, T., Sato, T., Hirose, K., Fukao, S., and Kato, S.: MU radar observations of the  
493 aspect sensitivity of backscattered VHF echo power in the troposphere and lower  
494 stratosphere, *Radio Sci.*, 21(6), 971–980, doi:10.1029/RS021i006p00971, 1986.

495 Tsuda, T., Kato, S., Yokoi, T., Inoue, T., Yamamoto, M., VanZandt, T. E., Fukao, S.,  
496 and Sato, T.: Gravity waves in the mesosphere observed with the middle and  
497 upper atmosphere radar, *Radio Sci.*, 25(5), 1005–1018, doi:  
498 10.1029/RS025i005p01005, 1990.

499 [Tsuda, T., VanZandt, T. E., and Saito H.: Zenith-angle dependence of VHF specular](#)  
500 [reflection echoes in the lower atmosphere, J. Atmos. Sol. Terr. Phys., 59,](#)  
501 [761–775, 1997.](#)

502 Vincent, R. A., Kovalam, S., Fritts, D. C., and Isler, J. R.: Longterm MF radar  
503 observations of solar tides in the low-latitude mesosphere: Interannual variability

504 and comparisons with the GSWM, *J. Geophys. Res.*, 103, 8667–8683, 1998.

505 [Waldteufel, P. and Corbin, H.: On the Analysis of Single-Doppler Radar Data, \*J. Appl.\*](#)

506 [\*Meteorol.\*, 18, 532–542, 1979.](#)

507 Wang, C.: New Chains of Space Weather Monitoring Stations in China. *Space*

508 *Weather*, 8: n/a. doi: 10.1029/2010SW000603, 2010.

509 Woodman, R. F., and Guille, A.: Radar Observations of Winds and Turbulence in the

510 Stratosphere and Mesosphere, *J. Atmos. Sci.*, 31, 493–505, 1974.

511 Woodman, R. F., and Chu, Y.-H.: Aspect sensitivity measurements of VHF backscatter

512 made with the Chung-Li radar: Plausible mechanisms, *Radio Sci.*, 24(2),

513 113–125, doi:10.1029/RS024i002p00113, 1989.

514 World Meteorological Organization (WMO): Measurements of upper air temperature,

515 pressure, and humidity, in *Guide to Meteorological Instruments and Methods of*

516 *Observation*, 6th ed., WMO I.12–1–I.12–32, Geneva, 1996.

517 Yamamoto, M., Fukao, S., Woodman, R. F., Ogawa, T., Tsuda, T., and Kato, S.:

518 Mid-Latitude E Region Field-Aligned Irregularities Observed With the MU

519 Radar, *J. Geophys. Res.*, 96(A9), 15,943–15,949, 1991.

520 Yamamoto, M. K., M. Oyamatsu, T. Horinouchi, H. Hashiguchi, and S. Fukao, High

521 time resolution determination of the tropical tropopause by the Equatorial

522 Atmosphere Radar, *Geophys. Res. Lett.*, 30(21), 2094, 2003.

523 Yang, P., Lu, D., Li, W., Wu, B., Fukao, S., Yamamoto, M., Tsuda, T., and Kato, S.:

524 Chaotic features of the atmospheric motions near the tropopause revealed by MU

525 radar observation: Case study, *Radio Sci.*, 25(5), 1065–1070, 1990.

526 Zhang S. D. and Yi, F.: A statistical study of gravity waves from radiosonde  
527 observations at Wuhan (30°N, 114°E), China. *Annales, Geophysicae*, 23,  
528 665-673, 2005.

529 Zhao Z. Y., Zhou, C., Qing, H. Y., Yang, G. B., Zhang, Y. N., Chen, G., Hu, Y. G.:  
530 Wuhan Atmosphere Radio Exploration (WARE) Radar: System Design and  
531 On-Line Winds Measurements, *Radio Sci.*, 48, 326–333, 2013.

532 Zink, F., Vincent, R. A., Murphy, E., and Cote, O.: Comparison of radar and in situ  
533 measurements of atmospheric turbulence, *J. Geophys. Res.*, 109, D11108, 2004.

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Table 1. WAER Radar Technical Parameter

Aspect	Specifications
<b>Radar system</b>	
Operating Frequency	53.8MHz ( $\lambda = 5.576\text{m}$ )
Power Synthesis	All solid state, Fully distributed
Peak Power	~172kW
Duty Cycle	Low mode 10% Medium mode 20% High mode 20%
<b>Antenna system</b>	
Antenna Array	24×24, active phased array
Antenna Type	Yagi aerial, 3 units, horizontal polarization
Normal Beam Width	$\leq 4.5^\circ$ half-power width, pencil beam
Voltage Standing Wave Ratio	$\leq 1.1$
Beam Direction	Five beams: vertical, off-zenith $0^\circ - 20^\circ$ by $1^\circ$
Antenna Operation Mode	Doppler Beam Swinging (DBS)
<b>Experimental Specifications</b>	
Pulse Width	1 $\mu\text{s}$ (Low mode) 32 $\mu\text{s}$ (Medium mode) 128 $\mu\text{s}$ (High mode)
Interpulse Period (IPP)	160 $\mu\text{s}$ (Low mode) 320 $\mu\text{s}$ (Medium mode) 1280 $\mu\text{s}$ (High mode)
No. of Coherent Integration	128 (Low mode) 64 (Medium mode) 8 (High mode)
No. of Incoherent Integration	10 (All modes)
No. FFT Points	256 (Low mode) 256 (Medium mode) 512 (High mode)
Time Resolution	1 minute

539 **Figure Captions**

540 **Figure 1.** (a) Power spectrum of four oblique beams. (b) Comparison of wind field  
541 observations between MST radar and rawinsonde for the altitude range of 3.2 to 16.9  
542 km at 16:00 LT on September 11, 2011: (left) wind speed and (right) wind direction.

543 **Figure 2. WARE observation of vertical wind velocity for 24 hours on October 2,**  
544 **2012.**

545 **Figure 23.** WARE observations of mesospheric wind: (a) the spectrum power of four  
546 oblique beams, and (b) the meridional and zonal wind estimated from 69 to 85 km.

547 **Figure 34.** (a)  $C_n^2$  profile (green line) estimated by radar echo power and  
548 temperature profile (blue line) estimated by rawinsonde. (b) WARE observations of  
549 the radar echo power from 3.2 km to 16.9 km for continuous 20 hours. The black  
550 dotted line denotes the height of radar tropopause.

551 **Figure 45.** Vertical profiles of (a) meridional, (c) zonal and (e) vertical wind  
552 disturbance components of the QM IGW observed at 06:05 LT on September 26,  
553 2011. (g) is the hodograph of the fitted meridional wind versus zonal wind. (b), (d), (f)  
554 and (h) are the same as (a), (c), (e) and (g), while the time is 13:05 LT on September  
555 26, 2011.

556 **Figure 56.** Observations of the variation of echo power as a function of beam zenith  
557 angle at six specified heights at 4.75 km, 6.09 km, 9.13 km, 13.13 km, 15.24 km and  
558 17.85 km by low and medium mode of WARE radar.

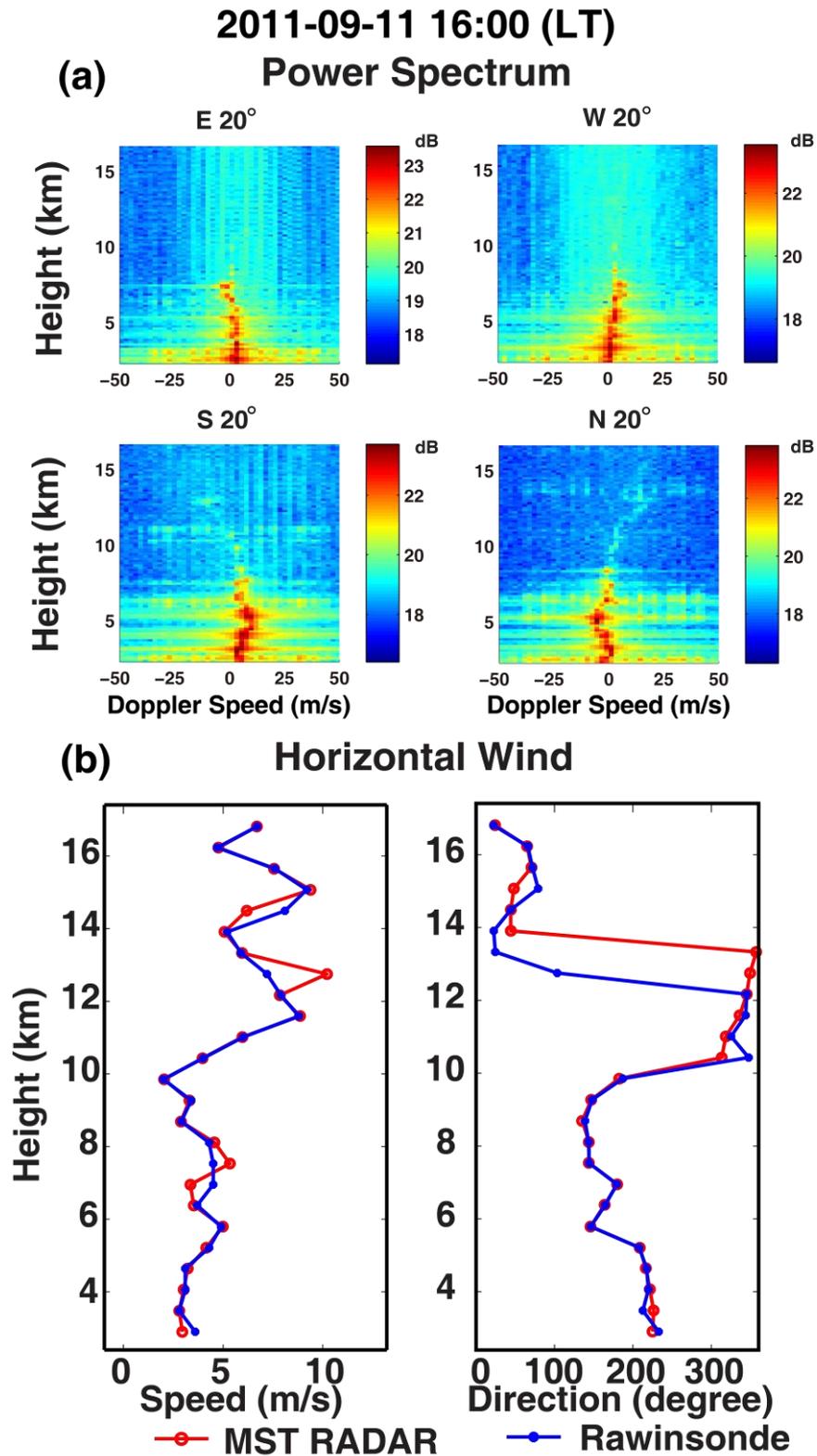
559 **Figure 67(a)** The raw time series for the meridional wind observed by WARE radar

560 from October 1 to October 12, 2011, with height resolution of 0.15 km and temporal  
561 resolution of 30 minutes. **(b)** Lomb-Scargle periodogram meridional wind  
562 disturbances. Only the values with confidence levels greater than 95% are shown.

563 **Figure 7.8.** Similar to Fig. 6, but for the zonal wind.

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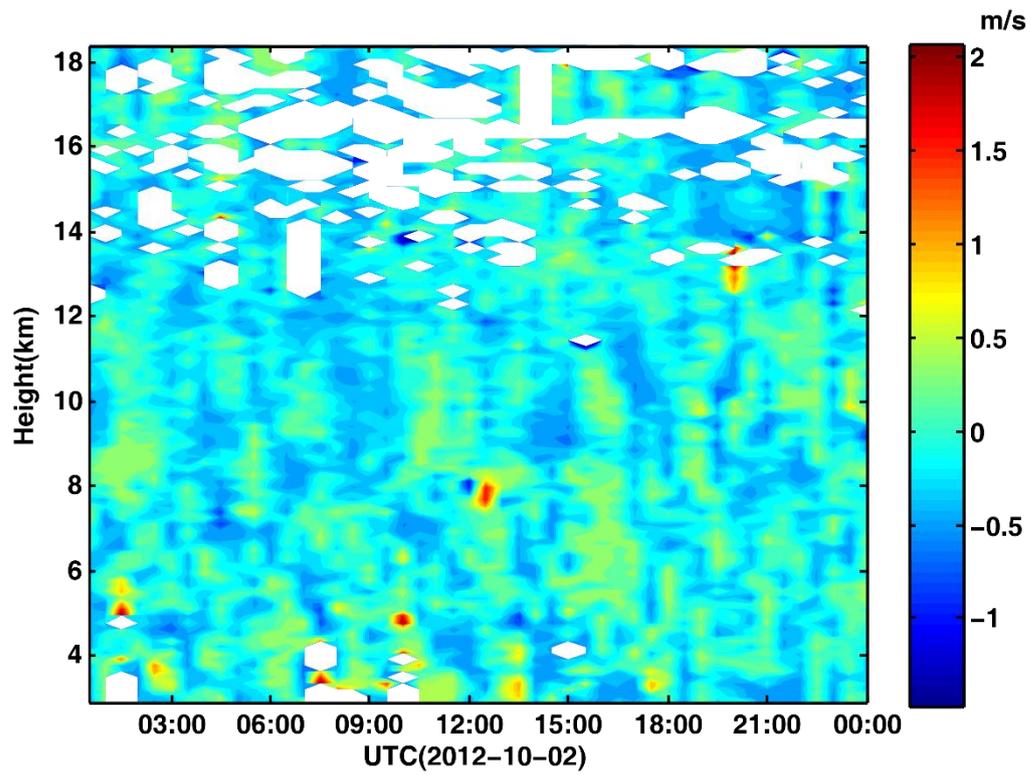


566

567 **Figure 1.** (a) Power spectrum of four oblique beams. (b) Comparison of wind field

568 observations between MST radar and rawinsonde for the altitude range of 3.2 to 16.9

569 km at 16:00 LT on September 11, 2011: (left) wind speed and (right) wind direction.



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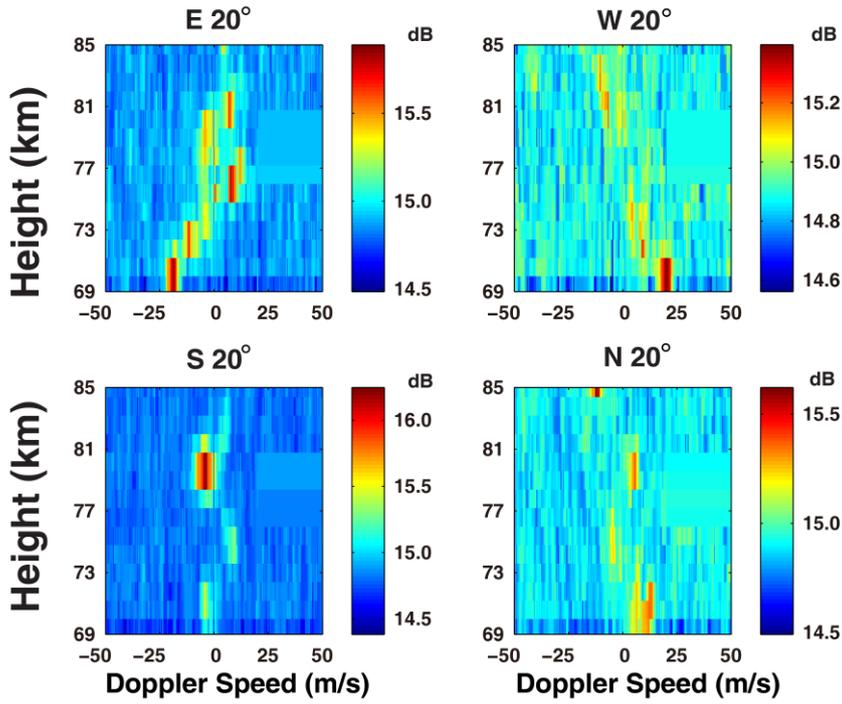
571 Figure 2. WARE observation of vertical wind velocity for 24 hours on October 2,

572 2012.

2011-3-15 13:44:24 (LT)

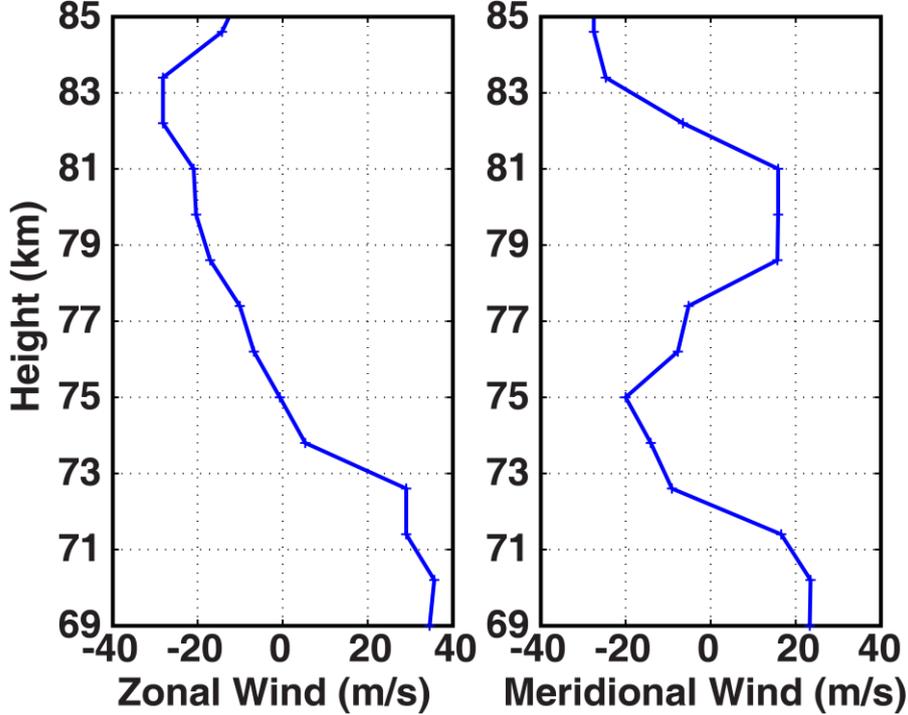
(a)

Power Spectrum



(b)

Horizontal Wind



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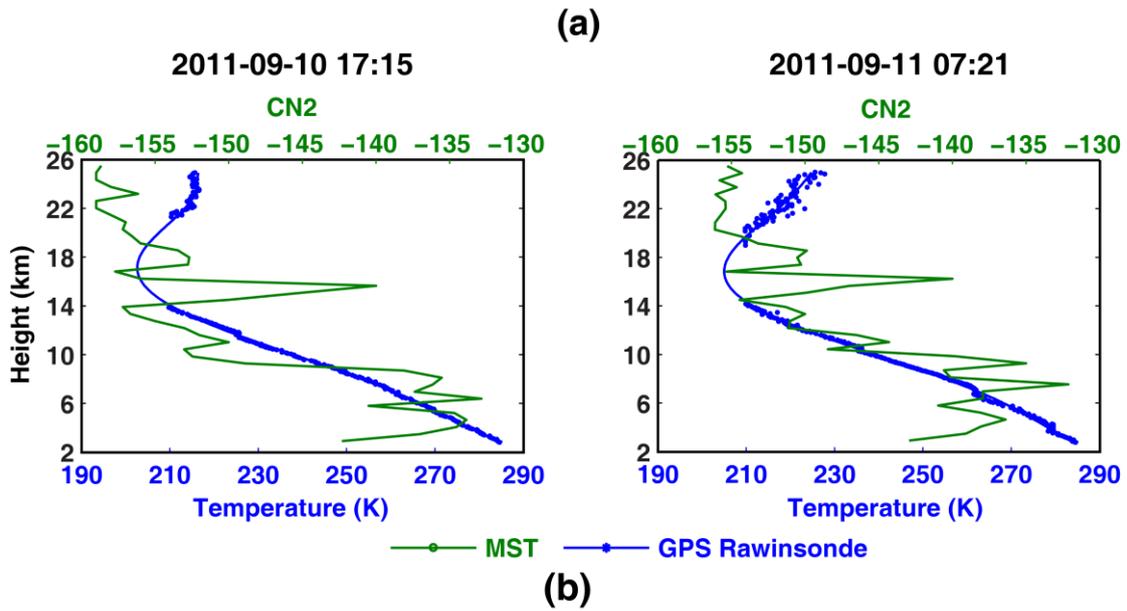
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**Figure 23.** WARE observations of mesospheric wind: (a) the spectrum power of four

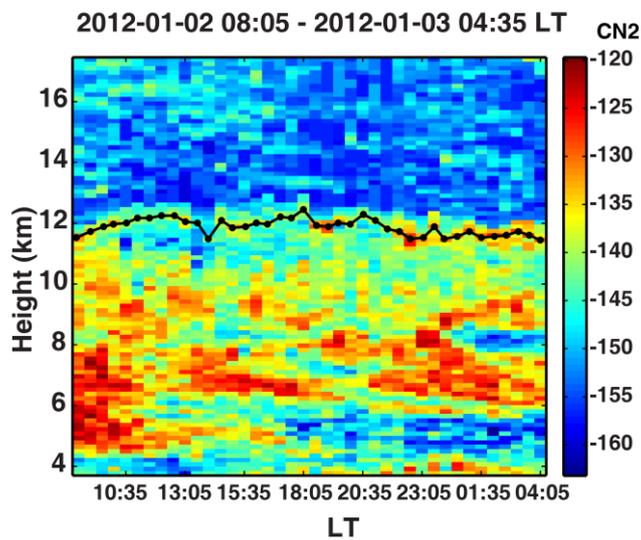
575

beams, and (b) the meridional and zonal wind estimated from 69 to 85 km.

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579 **Figure 34.** (a)  $C_n^2$  profile (green line) estimated by radar echo power and

580 temperature profile (blue line) estimated by rawinsonde. (b) WARE observations of

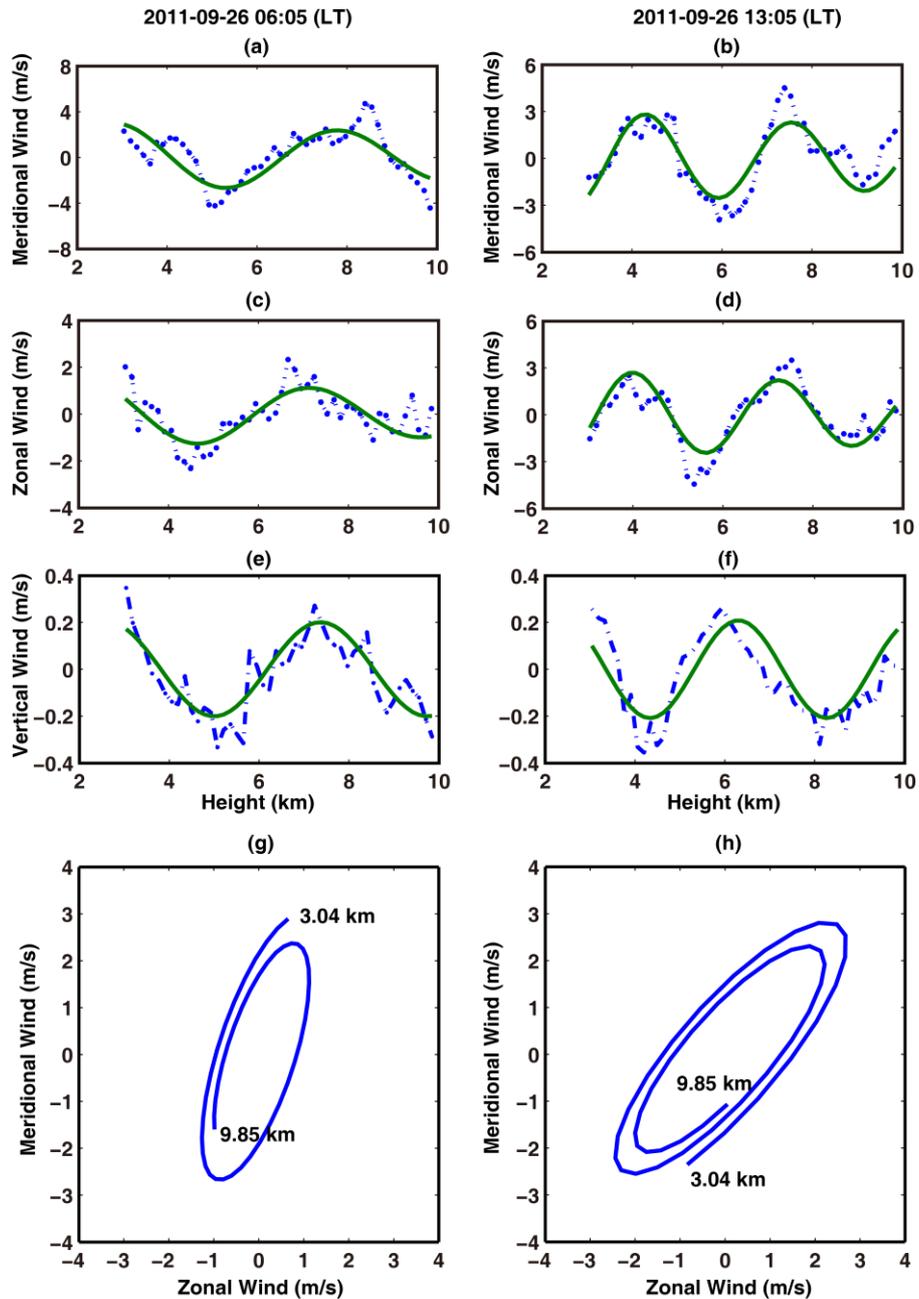
581 the radar echo power from 3.2 km to 16.9 km for continuous 20 hours. The black

582 dotted line denotes the height of radar tropopause.

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Figure 4

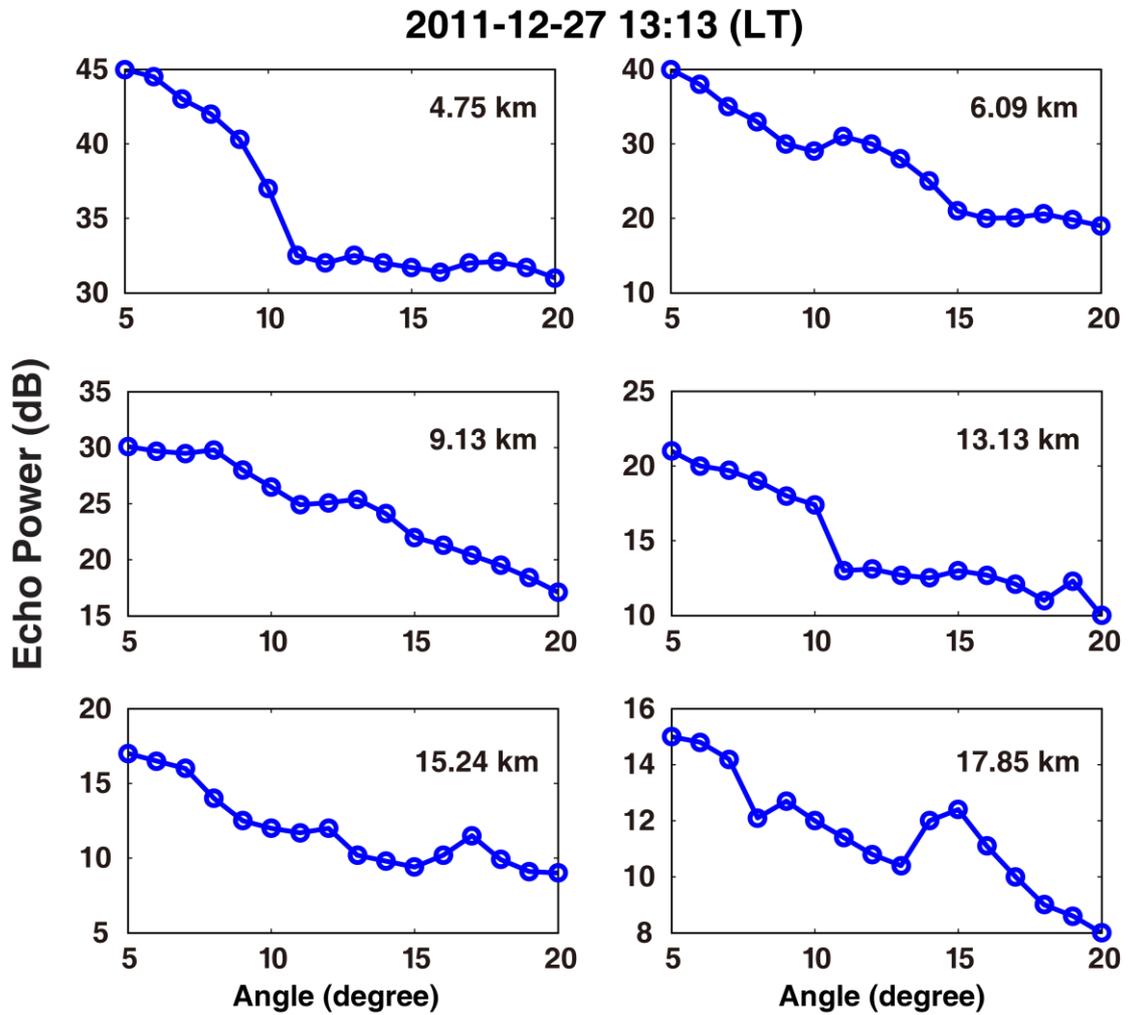


586

587 **Figure 45.** Vertical profiles of (a) meridional, (c) zonal and (e) vertical wind  
 588 disturbance components of the QM IGW observed at 06:05 LT on September 26,  
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 590 and (h) are the same as (a), (c), (e) and (g), while the time is 13:05 LT on September  
 591 26, 2011.

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Figure 5



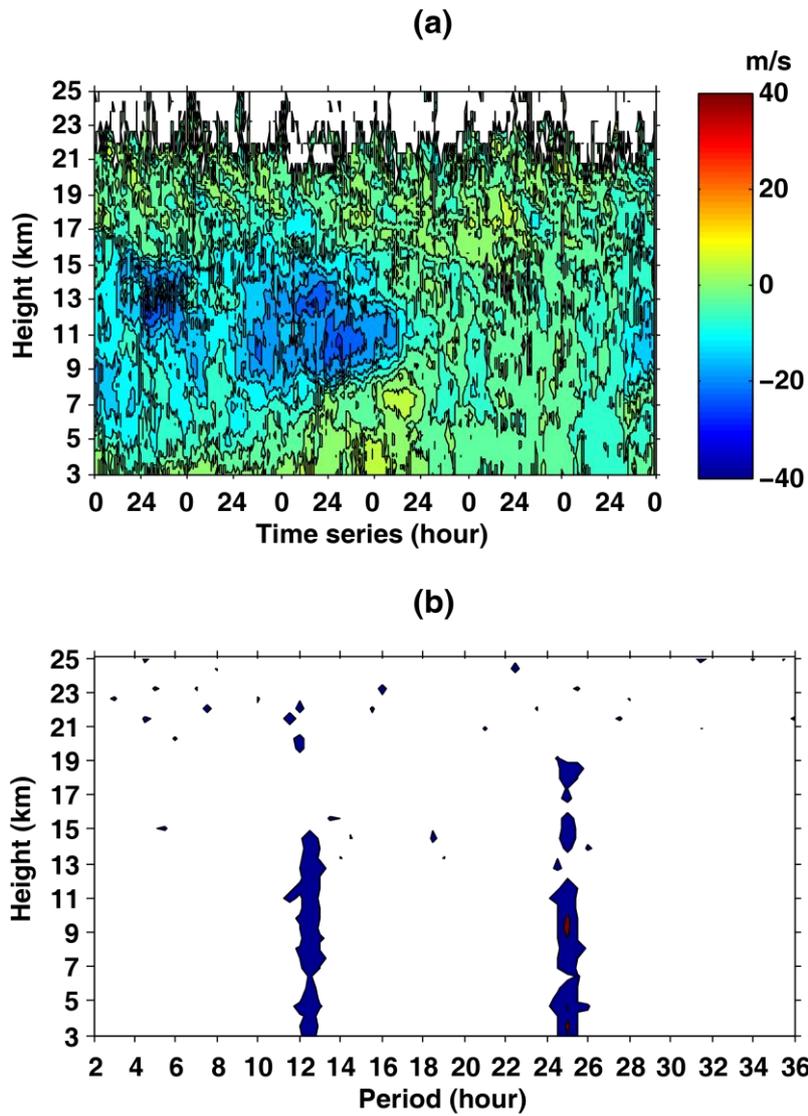
594

595 **Figure 56.** Observations of the variation of echo power as a function of beam zenith

596 angle at six specified heights at 4.75 km, 6.09 km, 9.13 km, 13.13 km, 15.24 km and

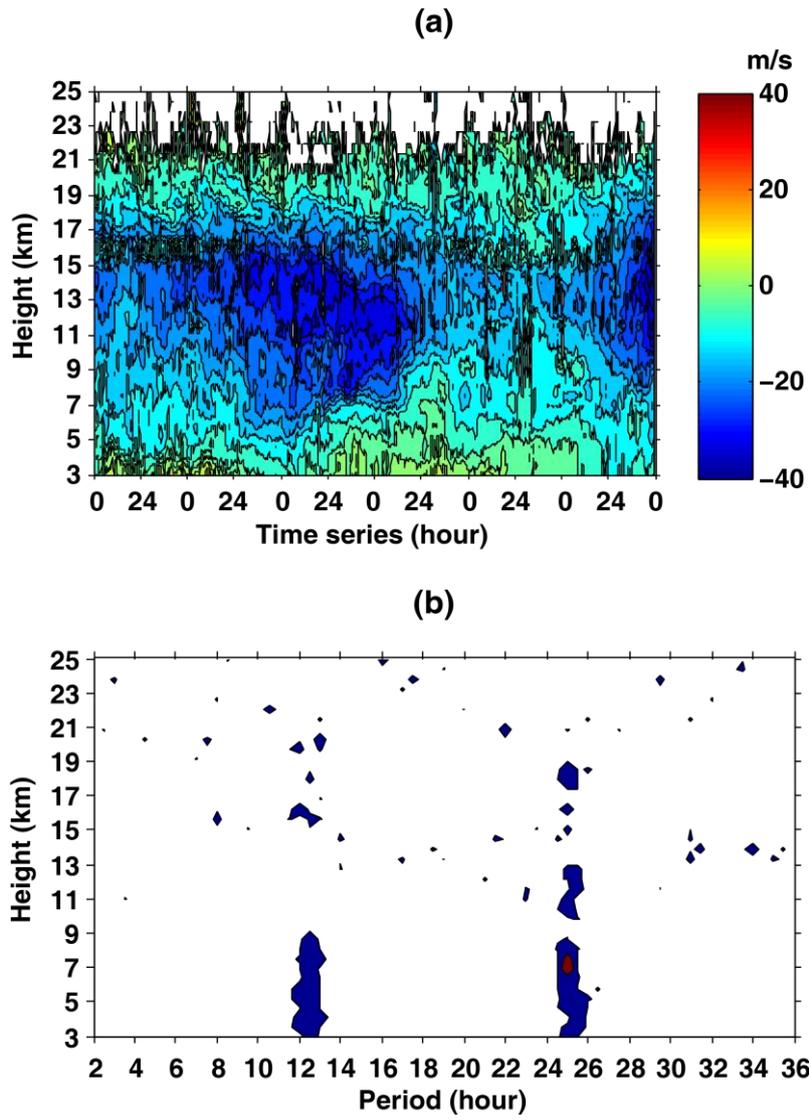
597 17.85 km by low and medium mode of WARE radar.

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600 **Figure 6-7** (a) The raw time series for the meridional wind observed by WARE radar  
 601 from October 1 to October 12, 2011, with height resolution of 0.15 km and temporal  
 602 resolution of 30 minutes. (b) Lomb-Scargle periodogram of meridional wind  
 603 disturbances. Only the values with confidence levels greater than 95% are shown.



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605

**Figure 78.** Similar to Fig. 6, but for the zonal wind.